

1 **High-resolution back-projection at regional distance: application to the Haiti**
2 **M7.0 earthquake and comparisons with finite source studies**

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8

9 **Abstract:**

10 A catastrophic Mw 7 earthquake ruptured on January 12th 2010 on a complex fault
11 system near Port-au-Prince, Haiti. Offshore rupture is suggested by aftershock
12 locations and marine geophysics studies but its extent remains difficult to define using
13 geodetic and teleseismic observations. Here we perform the multi-taper MUSIC
14 analysis, a high resolution array technique, at regional distance with recordings from
15 the Venezuela National Seismic Network to resolve high frequency (about 0.4 Hz)
16 aspects of the earthquake process. Our results indicate westward rupture with two
17 subevents, roughly 35 km apart. In comparison, a lower frequency joint finite source
18 inversion, with fault geometry based on new geologic and aftershock data, shows two
19 slip patches with centroids 21 km apart. Apparent source time functions obtained
20 from USArray further constrain the inter-subevent time delay, which implies a rupture
21 speed of 3.3 km/s. The tips of the slip zones coincide with the subevents imaged by
22 back-projections. The different subevent locations found by back-projection and by
23 source inversion suggest spatial complementarity between high and low frequency
24 source radiation, consistent with high frequency radiation originating from rupture
25 arrest phases at the edges of main slip areas. The GCMT solution and a geodetic-only

26 inversion have similar moment, indicating most of the moment release is captured by
27 geodetic observations without requiring additional rupture offshore. Our results
28 demonstrate the contribution of source imaging by back-projections of regional
29 seismic array data for earthquakes down to $M \approx 7$, especially when incomplete
30 coverage of seismic and geodetic data imply large uncertainties in source inversions.

31

32 **1. Introduction**

33

34 The M7 earthquake that hit Haiti in January 2010 was one of the most devastating
35 natural disasters of the last decades. The disproportionate damage caused by this
36 event and the prospect of future hazards in the region have prompted efforts to better
37 understand its rupture process and the tectonics of the Northeast Caribbean, in
38 particular the Enriquillo-Plantain-Garden Fault (EPGF) and the surrounding fault
39 systems. [Nettles and Hjorleifsdottir, 2010] found that a composite source model with
40 a strike-slip sub-event followed by a smaller thrust sub-event was consistent with
41 long-period teleseismic data. Using geodetic data, [Calais *et al.*, 2010; Hashimoto *et*
42 *al.*, 2011] attributed the slip to two asperities on a previously unmapped fault in the
43 Leogane delta, north from the EPGF. Based on the analysis of geodetic, geologic and
44 teleseismic observations [Hayes *et al.*, 2010] also inferred rupture primarily on the
45 blind Leogane fault, with a minor contribution on the EPGF. One of the outstanding
46 uncertainties about this event is the extent of its rupture in offshore areas that are
47 poorly constrained by on-land geodetic data. The possibility of coseismic slip offshore
48 is suggested by active deformation identified by a marine geophysical survey and by
49 the location of aftershocks derived from a combined temporary seismic network of 4
50 land and 21 ocean bottom stations (Haiti-OBS campaign) [Mercier de Lépinay *et al.*,

51 2011].

52 While finite fault inversions provide possible source models, they rely on a priori
53 information about the fault geometry, which is not readily available for the Haiti event
54 due to the geological complexity of the fault system and to the lack of surface rupture
55 [Prentice *et al.*, 2010] and strong motion recordings. In addition, the source inversion
56 problem suffers from limited resolution of the spatio-temporal rupture process due to
57 its low frequency band. Source imaging by back-projection of body waves recorded
58 by dense arrays allows to track the areas of strongest high frequency radiation [Ishii *et*
59 *al.*, 2005; Fletcher *et al.*, 2006; Vallee *et al.*, 2008]. This technique constrains the
60 spatio-temporal properties of the rupture (length, direction, speed, segmentation)
61 based solely on the phase of coherent seismic array signals. It does not rely on
62 detailed knowledge of Green's functions and fault geometry, on restrictive
63 parameterizations of the rupture kinematics, nor on additional smoothing. The high
64 frequency aspects of the rupture process imaged by array back-projection are
65 complementary to traditional finite source inversion models based on teleseismic and
66 geodetic data, which are instead sensitive to low frequencies and to the static field.

67 Conventionally, back-projection is applied to large seismic arrays at teleseismic
68 distances [Ishii *et al.*, 2005; Meng *et al.*, 2011] or to small aperture strong-motion
69 arrays at local distances [Spudich and Cranswick, 1984; Fletcher *et al.*, 2006].

70 Seismic arrays at regional distances can provide higher aperture to distance ratio and
71 thus higher resolution. [Vallee *et al.*, 2008] exploited surface waves recorded by a
72 regional array to study the 2001 Mw 7.8 Kunlun earthquake over a frequency band
73 (0.04 to 0.1 Hz) adequate to study very long ruptures, but too low to resolve smaller
74 earthquakes. In principle, higher frequency body waves carry higher resolution
75 information. However, the complexity of regional Pn waves has prevented

76 seismologists from fully exploiting this phase for source imaging. Earthquake source
77 studies using regional body wave phases have been mainly limited to inferring
78 macroscopic source properties from recordings at distances up to a few hundred km
79 [Zhu and Helmberger, 1996; Mendoza, 2005; Wei et al., 2009]. [Guilbert et al., 2005]
80 imaged the rupture propagation of a very large event, the 2004 Mw 9.0 Sumatra
81 earthquake, by array processing of body waves recorded by the CMAR-seismic array
82 at regional distance. Here, we further show that the relatively sustained character of
83 the Pn phase enables the application of high resolution array processing techniques on
84 moderate earthquakes ($M \approx 7$) to provide complementary constraints on rupture length
85 and locations of high-frequency source radiation. This capability can contribute to
86 rapid hazard and damage assessment for future earthquakes in the Caribbean region.
87 The rest of this article is organized as follows. In section 2 we describe the data
88 recorded by Venezuela National Seismic Network and argue for the need of high
89 resolution regional array analysis. In section 3 we present a high resolution array
90 source imaging technique adapted to earthquakes recorded at regional distance and
91 quantify the resolution of the technique through extensive synthetic tests. In section 4
92 we present our results of source imaging of the Haiti earthquake by regional array
93 analysis. In section 5 we integrate these results with independent analysis of apparent
94 source time functions and rupture speed based on USArray data and an improved
95 finite fault model based on teleseismic and geodetic data. In section 6 we discuss the
96 frequency dependent source properties, the possibility of offshore coseismic rupture
97 and the potential contribution of regional arrays to earthquake studies.

98

99 **2. VNSN data, pre-processing and selection**

100 Our array back-projection study of the 2010 Haiti event is based on data recorded at

101 regional distance by the Venezuela National Seismic Network (VNSN). The VNSN is
102 composed of 22 broad-band stations, oriented mainly East-West, located
103 approximately 9.5 degrees from Haiti in the perpendicular direction to the EPGF
104 strike (Fig.1). The VNSN has a privileged geometrical configuration and location to
105 study large earthquake ruptures in the Caribbean region.

106 Array processing at regional distance aims at estimating the azimuth of arrival and
107 relative timing of the seismic phases radiated from the strongest subevents of the
108 earthquake. The study of the 2010 Haiti earthquake remains challenging because of its
109 compact source size, shorter than 40 km according to previous studies [*Calais et al.*,
110 2010; *Hayes et al.*, 2010]. The resolution length scale along the fault that can be
111 achieved with the standard beamforming techniques is evaluated by the array
112 response function [*Rost and Thomas*, 2002]. Figure 2 shows the array response of the
113 VNSN and the USArray, the nearest array at teleseismic distance. The array responses
114 are back-projected into the hypocentral region of the 2010 Haiti earthquake based on
115 P travel times computed by the Tau-P toolkit and the IASP91 model. The array
116 response of VNSN is not isotropic. It has low resolution along the range direction (the
117 source-to-array direction) but adequate cross-range resolution (sub parallel to the fault
118 strike in this case). In regional array processing, the range resolution is considerably
119 poorer than the cross-range resolution, due to the small variability of the slowness of
120 Pn waves as a function of epicentral distance. Hence, hereafter we map the array
121 analysis results onto the fault plane by projecting them along the major axis of the
122 array response pattern. The along-strike resolution length achieved by the USArray is
123 approximately twice of that of the VNSN due to its more distant location and its more
124 unfavorable orientation with respect to the fault strike. We note that the array
125 response provides an ideal estimate of the array resolution. In practice, resolution is

126 further affected by waveform incoherence and interference, as discussed through
127 synthetic tests in section 3.2. Thus, to achieve adequate imaging of the Haiti
128 earthquake the back-projection requires data from an array at regional distance.
129 Moreover, since previous studies indicate that the rupture length of the Haiti
130 earthquake is as compact as 40 km, the analysis requires an array processing
131 technique that can achieve higher resolution than standard beamforming.

132

133 The Pn waveforms recorded by the VNSN are filtered between 0.2 and 0.7 Hz. The
134 low frequency cutoff is dictated by a time-frequency trade-off (see section 3.1). The
135 high frequency cutoff is determined by the coherency of the array data, which is
136 ultimately determined by the spacing between stations. The dominant frequency of the
137 Pn waves is about 0.4 Hz. We align the waveforms on their first arrival by multi-
138 channel alignment [*Vandecar and Crosson, 1990*] based on cross-correlation of 10
139 seconds long windows containing the first Pn arrivals. This procedure reduces the
140 effect of travel time errors due to uncertainties in the velocity structure e.g. [*Ishii et*
141 *al., 2005*].

142 Array processing techniques assume coherent signals across the array. We select a
143 subset of 13 stations with adequate waveform coherency by inspecting the array data
144 coherence matrix. This matrix is made of the coherence between all pairs of stations
145 computed during the multi-channel alignment procedure. The indices of the matrix are
146 reordered to group together the most mutually coherent stations (Fig. 3). We found a
147 modified k-nearest neighbors clustering algorithm
148 (www.eigenvector.com/MATLAB/Mac_Mfiles/corrmap.m) to be adequate for this
149 purpose. After inspection of the reordered matrix we select a subset of 13 stations that
150 consistently have correlation coefficients larger than 0.8 (Fig.3).

151

152 **3. High resolution multi-taper/MUSIC technique for regional arrays**

153

154 **3.1. Outline of the method and parameter settings**

155 Here, we apply at regional distance a high-resolution array analysis technique
156 originally developed for a teleseismic array study [Meng *et al.*, 2011]. The method
157 combines the Multiple Signal Classification (MUSIC) array processing technique
158 [Schmidt, 1986; Goldstein and Archuleta, 1991; Guilbert *et al.*, 2005] with multi-
159 taper cross-spectrum estimation [Thomson, 1982]. MUSIC was designed for high
160 resolution direction of arrival estimation of long and stationary signals [Krim and
161 Viberg, 1996]. The Pn phase is relatively stationary. The multi-taper method yields a
162 robust estimate of the data coherence matrix on relatively short time window and thus
163 improves significantly the temporal resolution of MUSIC. Here we only describe the
164 choice of processing parameters for this particular study. More details about the
165 method are described by [Meng *et al.*, 2011].

166

167 *Number of tapers in cross-spectrum estimation*

168 The multi-taper technique [Thomson, 1982] averages multiple, almost independent
169 cross-spectral estimates obtained after tapering the data by a sequence of orthogonal
170 functions with optimal temporal and spectral concentration. These Slepian tapers are
171 the set of functions of finite temporal duration T with maximum energy within the
172 band of frequencies lower than the bandwidth W . Given a time window duration T
173 and a frequency bandwidth W , the proper number of tapers is $2 \times T \times W - 1$. As a
174 compromise between useful averaging (large $T \times W$) and adequate temporal resolution

175 (short T) we use here 3 tapers, which implies $T \times W = 2$.

176

177 *Length of analysis window*

178 The MUSIC method requires narrow band estimates of cross-spectrum, i.e. the
179 bandwidth W (the frequency smearing of the spectral estimator) has to be small
180 compared to any frequency of analysis, $W \ll f$. This implies a trade-off between
181 temporal resolution and frequency localization: $T \times f \gg 2$ for 3 tapers. As a
182 compromise here we set $T = 30$ s, which satisfies $T \times f > 6$ for frequencies higher than 0.2
183 Hz. At the dominant frequency of our data, $f \approx 0.4$ Hz, we get $T \times f \approx 12$.

184

185

186 *Number of principal eigenvalues*

187 In MUSIC the signal subspace is defined as the subspace spanned by the distinctly
188 largest eigenvalues of the data covariance matrix. Its complement defines the noise
189 subspace. For stationary signals, the dimension N of the signal subspace is equal to
190 the number of signal sources contributing to the data time window. In the case of
191 transient seismic signals analyzed with relatively short time windows, each
192 covariance matrix estimated with a separate taper tends to have only one non-zero
193 eigenvalue and the associated eigenvectors tend to be independent. Hence, the rank of
194 the covariance matrix obtained by linear combination of the multiple taper estimates
195 tends to be equal to the number of tapers. In particular, the number of significant
196 eigenvalues is at most three when using three tapers. Fig. 4 shows results of the
197 MUSIC analysis of the 2010 Haiti earthquake at 0.4 Hz, the dominant frequency of
198 the seismograms, with N ranging from 1 to 4. The MUSIC pseudo-spectrum, a
199 measure of the orthogonality between the array steering vector associated to a

200 candidate source location and the noise sub-space (for details see [Meng *et al.*, 2011]),
201 is plotted as a function of along-strike position with respect to the hypocenter and as a
202 function of end time of the sliding windows. The MUSIC pseudo-spectrum is
203 normalized by its maximum in each time window. Its amplitude depends on signal
204 power, coherence and interference but it does not directly represent any physical
205 quantity. Nevertheless, the sharp peaks of the MUSIC pseudo-spectrum indicate the
206 location of signal sources. In the MUSIC analysis of the Haiti earthquake, two main
207 subevents (black dots in Fig. 4b) are visible and their estimated locations are
208 independent of the choice of N . However, when N is too small ($N=1$) the analysis
209 fails to resolve two simultaneous sources, it only indicates the strongest source in each
210 time window. The results with $N=2$ or 3 show stable features of the rupture process
211 including simultaneous multiple sources. When N is larger than the number of non-
212 zero eigenvalues (e.g. $N=4$), the signal subspace is contaminated by noise and the
213 resulting image is not stable. While the spatial location of the subevents is
214 independent of the choice of N , their estimated timing is not. Because we have not
215 developed yet a complete understanding of the effect of N on the timing of subevents,
216 an objective rationale for the choice of N is not available at this time. Here we choose
217 $N=2$ because it yields the most clear image of the westward rupture front and the
218 resulting rupture speed is within the usual range, consistent with independent
219 observations available for this earthquake, as described in later sections. We note,
220 however, that the temporal details of the back projection images, which depend on N ,
221 are not used as additional constraints in our analysis of this earthquake.

222

223 **3.2. Resolution tests**

224 *Synthetic test for an ideally coherent linear array*

225 Fig. 5 compares the resolution of four techniques on synthetic transient signals:
226 beamforming, cubic root stacking (a popular modification of beamforming, e.g. [Rost
227 and Thomas, 2002]), correlation stacking and MUSIC. In correlation stacking the
228 normalized cross-correlation coefficients are beamformed instead of the waveforms
229 [Fletcher et al., 2006] to improve robustness against scattering, multipathing and
230 contamination by coda waves [Borcea et al., 2005]. The color scale in the images by
231 correlation stacking and cubic root stacking indicate beamformed cross-correlation
232 coefficients and beamformed cubic root of the signal amplitude, respectively,
233 evaluated or integrated over a sliding time window of duration 30 s The resolution of
234 a method is defined as its ability to separate closely spaced sources. We consider two
235 identical Pn plane waves with dominant frequency of 0.3 Hz impinging
236 simultaneously but with different azimuth on a linear array of 21 sensors regularly
237 spaced at half wavelength. Gaussian white noise is added with a signal to noise ratio
238 of 10 dB. The results as a function of the relative azimuth between the two waves
239 (Fig. 5) show that, the minimum azimuthal separation resolvable by beamforming,
240 cubic root stacking, correlation stacking and MUSIC is approximately 8, 8, 6 and 3
241 degrees, respectively. This shows that, under perfect waveform coherency, the
242 azimuthal resolution, and hence the spatial resolution in the cross-range direction,
243 achieved by our multitaper MUSIC method achieves azimuthal resolution that
244 outperforms the other methods by at least a factor of two.

245

246 *Synthetic test for the VNSN array geometry*

247 Fig. 6 shows further synthetic tests to understand the performance and potential bias
248 of the back projection of the Haiti earthquake source using the VNSN data. We
249 compute full wave field synthetics (Fig.7a) using the SPECFEM3D package [Tromp

250 *et al.*, 2008] with the 3D crustal velocity model Crust2.0 [*Bassin et al.*, 2000]. Our
251 simulations are accurate up to 0.33 Hz. Given the uncertainties in the velocity model,
252 generating higher frequency synthetics is not warranted. Inspired by the final result of
253 the back-projection, we test scenarios with a source containing two asperities. The
254 first one is located at the hypocenter of the Haiti earthquake. The second one is
255 located westward along the hypothetical Leogane fault trace [*Hayes et al.*, 2010]. We
256 consider either the same slip amplitude for both subevents or a 2:1 ratio based on the
257 source time functions estimated at the USArray (see section 5.1). We explore inter-
258 subevent distances up to 50km, with rupture time delays consistent with a rupture
259 speed of 3 km/s. For both asperities we assume source parameters from the CMT
260 solution (focal mechanism 251/70/28, strike/dip/rake, and centroid depth 12 km) and
261 a Brune source time function with duration of 3 seconds. In Fig.6, the left two
262 columns show examples of cubic root stacking, correlation stacking and MUSIC
263 pseudo-spectra. These quantities are back-projected onto the source area based on P
264 travel times computed by the Tau-P toolkit and the IASP91 model.

265

266 We found that all the back-projection methods are modulated by interference between
267 subevents. This introduces an amplitude and location fluctuation that depends
268 periodically on the product of inter-subevent time delay and frequency of analysis.

269 Fig. 8 shows two examples of interference in the cubic root stacking analysis at 0.3
270 Hz, with sources separated by 15 and 40 km respectively. The time delay between the
271 wave arrivals from the two asperities varies across the array. Destructive interference
272 occurs at stations where the waves from the two asperities are out of phase with
273 respect to the dominant period (the seismograms of some stations are deficient in low
274 frequencies in Fig. 8). This issue might be less severe in practice, since the synthetics

275 are computed from two point sources. In reality, the broader slip region reduces the
276 simultaneous destructive interference.

277 As expected from the resolution test, MUSIC achieves a smaller resolvable distance
278 than the other two techniques. This difference is particularly prominent when we set a
279 more realistic 2:1 amplitude ratio for the two sources. The left two columns in Figure
280 6 show the two snapshots taken at the beginning of each source ($t=0s$ and $t=10s$) when
281 $d=30$ km. The MUSIC pseudo-spectrum clearly gives two peaks in the second
282 snapshot, while the other two techniques fail to resolve them. Note that the absolute
283 amplitude of the MUSIC pseudo-spectrum, which is modulated by interference
284 patterns, does not directly represent the signal power and is not used in the analysis.

285 The resolution limit for MUSIC in these synthetic tests is approximately 25 km with
286 an uncertainty of about 5 km for each sub-source. Considering the uncertainty is
287 mainly introduced by the interference effect, which is less prominent in practice, 5 km
288 is regarded as the upper bound of the relative location uncertainty in our analysis.

289

290 **4. Results of array analysis of the 2010 Haiti earthquake at regional distance**

291

292 We first illustrate the accuracy of our method by an analysis of the largest aftershock,
293 a M 5.9 (NEIC) earthquake on January 20th 2010 (waveform shown in Fig.7b). The
294 back-projection of the MUSIC pseudo-spectrum into the source region is shown in
295 map view in the first two columns of Fig. 9d for two windows ending at $t=15$ s and 30
296 s after the first arrival, respectively. The warmest color indicates the location of the
297 source of the waves arriving at the VNSN within these time windows. As explained in
298 section 2, the back-projected contours are very elongated in the source-array
299 direction. Our analysis provides accurate source location only in the perpendicular

300 direction, which is sub-parallel to the EPGF strike. We hence extract the maximum of
301 the MUSIC pseudo-spectrum along lines parallel to the source-array direction, then
302 project these maxima onto the EPGF trace. Note that this aftershock has a slightly
303 different strike (N290°) than the EPGF (N255°) but this difference is minor for this
304 magnitude. The right column of Fig. 9 shows this quantity as a function of epicentral
305 distance along the EPGF strike and of final time of the sliding window. Fig.9-d shows
306 the aftershock is properly imaged as a compact source of 5 km size near its
307 hypocenter, comparable to the typical rupture size of M6 earthquakes [*Wells and*
308 *Coppersmith, 1994*]. This suggests little spatial smearing in our analysis at the 5 km
309 scale and warrants imaging of the mainshock, which is possibly composed of multiple
310 subevents.

311 Fig. 9c shows the results of our application of MUSIC to the VNSN recordings of the
312 2010 Haiti mainshock. We also applied two other popular array techniques, cubic-root
313 stack and correlation stack with the same stations set and frequency band, shown in
314 Figs. 9a and 9b respectively. Although all these array analysis techniques
315 unambiguously indicate two prominent high-frequency subevents, our multitaper-
316 MUSIC algorithm provides the images with the sharpest contrast between the
317 subevents and the adjacent areas, as expected from our synthetic tests. Note that the
318 peak locations of the MUSIC imaging are off the fault trace, to the South, due to
319 limited resolution along the source-array direction at regional distance, which
320 highlights the importance in regional array processing of a projection along the range
321 direction onto the fault trace. The first subevent is located approximately 7 km east
322 from the hypocenter. The second subevent is approximately 35 km further west
323 (28km west from the hypocenter). This complexity is absent in the results of our
324 analysis of the M6.1 aftershock (Fig.9-d) and hence is not due to path effects or other

325 phases. The temporal evolution is better imaged by array processing at the peak
326 frequency (0.4 Hz). In Fig.4-b, bilateral propagation is observed starting around $t=10$
327 s, when the last 10 s of the 30 s window contain the signal. The eastward front runs
328 for about 10 km in about 3 s. The westward front shows a weaker MUSIC pseudo-
329 spectra but its propagation can be consistently tracked over a distance of ~ 35 km and
330 a duration of about 10 s (until $t\approx 20$ s). The rupture speed of both fronts is similar and
331 consistent with usual sub-shear rupture speeds, given the shear wave speed of 3.65
332 km/s at 10 km depth in the Haiti region [Bassin *et al.*, 2000]. The observed features
333 are compatible with the USArray source time functions and finite fault inversions
334 described in the next section.

335

336 **5. Comparison to independent observations**

337

338 Our study is the first example of earthquake source imaging at regional distance with
339 Pn waves for earthquakes of magnitude as low as 7. To build confidence in our
340 results, we compare them here to other available observations. The complementarity
341 between the different source analysis techniques provides an integrated view of the
342 2010 Haiti earthquake rupture process.

343

344 **5.1. Source time functions at USArray and inferred rupture speed**

345

346 The temporal separation between the two asperities is independently estimated from
347 apparent source time functions obtained through deconvolution of teleseismic data by
348 theoretical Green's functions (Fig.10) [Chu *et al.*, 2009]. We selected USArray
349 stations with epicentral distances greater than 30 degrees to avoid complexities of the

350 P waveform due to triplication in the upper mantle. We computed synthetic Green's
351 function using the generalized ray theory [Helmberger, 1983], the IASPEI91 1D
352 global velocity model [Kennett and Engdahl, 1991] with the upper 30 km adjusted to
353 a 4-layer Haiti regional crust model [Bassin et al., 2000], and the CMT solution. The
354 vertical component P waveforms and synthetic Green's functions are filtered from
355 0.01 to 1Hz and deconvolved in the time domain with the iterative technique of
356 [Kikuchi and Kanamori, 1982] and with the non-negative least-squares algorithm.
357 Both deconvolution techniques yield consistent results. To estimate the temporal
358 separation between the subevents and its uncertainty, we considered 1000 stacks of
359 randomly selected subsets containing 90% of the USArray stations. We measured the
360 delay between the two prominent peaks in the stacked apparent source time functions.
361 The resulting delays have a Gaussian distribution with mean $t_1=5$ s and standard
362 deviation 0.02 s (Fig.10). We associate these to the two subevents found by our array
363 analysis. The deconvolution analysis also indicates that the second subevent accounts
364 for approximately one third of the total moment release, consistent with the multiple
365 CMT solution of [Nettles and Hjorleifsdottir, 2010]. Their analysis also indicates that
366 the second subevent has a thrust mechanism, in contrast to the near strike-slip
367 mechanism of the first subevent. Nevertheless, at the azimuth of the USArray both
368 mechanisms produce similar amplitudes and our choice to use the CMT solution to
369 generate the Green's functions does not introduce a significant bias. The rupture
370 speed can be reliably constrained by the apparent subevent delay of $t_1=5$ s and by the
371 spatial subevent separation of 21 km between the two subevents centroids provided by
372 our geodetic-only slip inversion (Figure S5, introduced in section 5.2). Considering a
373 bilateral rupture, the real inter-subevent time delay is $t_0 = t_1 + \cos\theta(L_w - L_e)/c$,
374 where θ is the angle between the fault strike and the source-array direction (40

375 degrees), c is the apparent P-wave speed (12 km/s at epicentral distance of 40
376 degrees), L_w and L_e are the distances of the centroid of the westward and eastward
377 subevents, respectively ($L_w = 21$ km and $L_e = 0$ km). This yields a true inter-subevent
378 time of 6.3 seconds. The implied rupture speed, 3.3 km/s, is within the usual sub-
379 shear range. For reference the shear wave speed is 3.65 km/s at 10 km depth in the
380 Haiti region [Bassin *et al.*, 2000]. Although our regional array back-projection
381 analysis currently has large uncertainties on the rupture timing (see section 3.1), our
382 back-projection result at the dominant frequency, 0.4 Hz (Fig. 4-b), is consistent with
383 this sub-shear rupture speed.

384

385 **5.2. Improved finite fault model**

386

387 To evaluate how the results of the back-projection and USArray analysis of the source
388 time function relate to other data sets and integrate in the global description of the
389 rupture, we perform a joint source inversion based on teleseismic body waves and the
390 various geodetic data available.

391 As detailed in previous studies [Calais *et al.*, 2010; Hayes *et al.*, 2010; Hashimoto *et*
392 *al.*, 2011], the campaign GPS measurements and InSAR data provide a very good
393 coverage of the Leogane delta where the rupture started and indicate that the
394 deformation in this area is best explained by a fault segment dipping NNW. In
395 agreement with those studies, we choose a fault strike of 254° and a dip angle of 55° .
396 This orientation of the fault plane is also compatible with the alignment of aftershocks
397 along a $N75^\circ$ azimuth [Mercier de Lépinay *et al.*, 2011]. InSAR fringes (ALOS
398 ascending track 137) and field measurements of the coastal vertical uplift in the area
399 of the town of Petit Goave [Hayes *et al.*, 2010] suggest the earthquake also extended

400 west and offshore of the Leogane delta. This hypothesis is supported by the aftershock
401 catalog obtained from the Haiti-OBS local network of terrestrial and ocean-bottom
402 seismic stations. In contrast with previous models, we do not cover this offshore area
403 by simply extending the preferred Leogane fault geometry (i.e. dipping NNW) to the
404 west, but refine the model with a distinct second fault segment dipping 45° to the
405 NNE (strike $N275^{\circ}$). This NNW to NNE rotation in the fault strike is supported by the
406 moment tensors of the aftershocks [*Nettles and Hjorleifsdottir, 2010*] and by the
407 identification of bathymetric features with a similar orientation, such as the Trois
408 Baies fault and the Transhaitian belt [*Mercier de Lépinay et al., 2011*].

409 The location of the NEIC epicenter seems incompatible with our first fault segment as
410 it would imply a very superficial rupture initiation (~ 2 km), and would not be
411 compatible with the NEIC hypocentral depth (13 km). Moving the hypocenter in the
412 northeast direction as suggested by the higher resolution Haiti-OBS aftershocks
413 catalog (when compared to the NEIC catalog) brings the rupture initiation closer to
414 the zone of high slip, as we would expect from the sharp onset of the rupture, and the
415 hand-picked P wave onset (Fig. S1). The applied shift of a few kilometers is
416 comparable to that applied by [*Mercier de Lépinay et al., 2011*] in their finite fault
417 inversion. To match this new epicenter location ($72.54^{\circ}W$, $18.46^{\circ}N$) to our fault
418 geometry, the hypocenter is fixed at a depth of 6 km.

419 The finite fault inversion is based on a Monte-Carlo type algorithm [*Ji et al., 2002*]
420 and uses teleseismic data (21 P waves and 13 SH waves bandpass filtered between 2
421 and 100 s), campaign GPS (data processing detailed in [*Calais et al., 2010*]) and four
422 InSAR images (data selection and processing detailed in [*Hayes et al., 2010*]). We
423 allow the rupture speed to vary from 3.0 to 3.6 km/s, that is ± 0.3 km/s around the

424 rupture velocity inferred from the USArray analysis. We do not apply any
425 minimization criteria on the seismic moment.

426 We obtain a slip model composed of two high slip patches, a first one centered on the
427 epicenter and a second one with peak slip amplitude 21 km further west, starting 6 s
428 after the onset of rupture (Fig.11). The data misfit in our updated model (Table S1) is
429 comparable to previous studies [*Calais et al.*, 2010; *Hayes et al.*, 2010; *Hashimoto et*
430 *al.*, 2011]. While the fit of the InSAR data is excellent (Fig. S2 and S3), the fit of the
431 GPS data is of moderate quality (Fig.11): this is mainly due to the large misfit at
432 station DFRT which, despite having one of the largest measured displacements (0.7 m
433 measured in the center of the deformation area), cannot be adjusted either in azimuth
434 or amplitude. At the location of station DFRT, ascending and descending InSAR
435 tracks are too decorrelated to do a comparison of the 3D displacement. However, the
436 smooth 3D pattern of deformation inferred from InSAR in the surrounding areas
437 [*Hayes et al.*, 2011] suggests that motion at DFRT is indeed affected by localized
438 deformation. *Calais et al.* [2010] who also used InSAR and GPS data could not fit
439 properly station DRFT. The vertical deformation along the coast was measured using
440 coral data [*Hayes et al.*, 2010]. Since most of those measurements are within the area
441 where InSAR coverage allows to estimate the vertical motion (overlap of ascending
442 and descending tracks) we did not include them in the inversion and only made sure
443 that they are compatible *a posteriori*. This is true for all points (Fig. 11) except one
444 subsidence measurement west of the surveyed area. This mismatch suggests that this
445 point indeed represents a local effect, perhaps a local slump as described by [*McHugh*
446 *et al.*, 2011] and [*Mercier de Lépinay et al.*, 2011] in the offshore survey of this area.
447 Also, this subsidence very close from uplifted points reflects a change which is
448 probably too localized to be properly reproduced by our simplified fault model.

449

450 Slip reaches a maximum of 10 meters (Fig. S4) at the center of the first patch (7 m
451 for the second patch) which concentrates $\frac{2}{3}$ of the total moment ($3 \cdot 10^{+19}$ N.m for the
452 first patch, $1.5 \cdot 10^{+19}$ N.m for the second). This result is similar to our analysis of the
453 USArray data. In term of bulk parameters, we also find that the moment tensor of our
454 solution (Fig. 11) and its seismic moment ($4.5 \cdot 10^{+19}$ N.m) are similar to the GCMT
455 analysis [Nettles and Hjorleifsdottir, 2010]. The teleseismic data bound the duration
456 of the source time function to less than 15 s and indicate that the first patch of slip was
457 more impulsive than the second and smaller slip episode. The oblique rake along the
458 rupture is consistent with the transcompressional regime in this region [Mercier de
459 Lepinay et al., 2011] but doesn't support the hypothesis of a fully partitioned fault
460 system where the EPGF would absorb all the strike-slip component [Dixon et al.,
461 1998; Calais et al., 2002]

462

463 **6. Discussion**

464

465

466

467 **6.1. Spatial complementarity between high and low frequency source properties**

468

469 The primary results from the array analysis, the finite fault inversion and the source
470 time function analysis suggest that the rupture propagates mostly westward breaking
471 two main slip patches. In the low frequency view from finite fault inversion the first
472 asperity is located near the epicenter and the second one is 21 km west and

473 contributes to one third of the total seismic moment. On the other hand, the high
474 frequency radiations suggest a distance of 35 km between the two asperities.

475 Since the back-projection provides subevents locations relative to the hypocenter, the
476 choice of epicenter is essential in interpreting the result of the array processing.

477 Assuming the epicenter given by NEIC, the subevent locations are consistent with the
478 spatial pattern of the aftershocks (NEIC catalog), which cluster into two groups. The
479 eastern group is close to the mainshock epicenter and the second group is located 30
480 to 40 km west. If we assume the NEIC hypocenter, the western subevent is well
481 beyond the high slip patches derived from the finite fault model, implying possible
482 slip further west offshore, unconstrained by geodesy. However, if we account for the
483 NEIC catalog bias inferred from the Haiti-OBS aftershock catalog in Fig. 11 [*Mercier*
484 *de Lépinay et al.*, 2011], the two subevents coincide with the terminal edges of the
485 large slip areas. This difference is similar to the spatial complementarity between high
486 and low frequency source properties inferred for several other earthquakes, including
487 the 1989 Loma Prieta earthquake, the 1993 Koshi-Oki earthquake [*Nakahara*, 2008]
488 and the 2011 Tohoku-Oki earthquake [*Ide et al.*, 2011; *Meng et al.*, 2011; *Simons et*
489 *al.*, 2011]. This complementarity can be interpreted as high frequency radiation
490 generated by the stopping phases, associated with abrupt rupture speed reduction at
491 the edge of the slip area. This mechanism of high-frequency radiation was first
492 described in circular crack models [*Madariaga*, 1977; 1983] and was explained for
493 general rupture front geometries through isochrone theory [*Bernard and Madariaga*,
494 1984; *Spudich and Frazer*, 1984]. This interpretation is also consistent with the
495 aftershock locations clustering at the regions of stress concentration at the edges of
496 the coseismic rupture. Note that the high frequency radiation is absent at the western

497 end of the first asperity, implying a smoother stopping or rupture transition without
498 stopping between the two asperities.

499

500 **6.2. Extent of the offshore rupture**

501 One aspect of the 2010 Haiti earthquake which still remains to be elucidated is the
502 extent to which the rupture propagated offshore in areas that are not resolved by
503 geodetic data. While there is no direct offshore geodetic measurements, InSAR and
504 GPS data cannot be properly explained by on-land faulting alone. GPS station LEOG
505 strongly points offshore and west of the Leogane delta (Fig. 11) despite the existence
506 of an asperity east of that point (this asperity is directly visible in the descending
507 InSAR track 447). The dense InSAR fringes of ascending track 137 (Fig. S3) surround
508 the coastline pointing to deformation in the center of the bay. All slip models based on
509 either InSAR or GPS data found at least partial offshore faulting [*Calais et al.*, 2010;
510 *Hayes et al.*, 2010; *Hashimoto et al.*, 2011; *Mercier de Lépinay et al.*, 2011], but the
511 western extent of off-shore slip is not constrained by these data.
512 Recent observations in the offshore region are suggestive of rupture extending further
513 west than in our kinematic source model. Marine seismic reflection profiles indicate
514 the existence of a large scale active anticline associated with the Trois Baies fault and
515 the Transhaitian belt [*Mercier de Lépinay et al.*, 2011]. The aftershock locations
516 derived from the Haiti-OBS campaign are offset by about 20 km NE relative to the
517 NEIC catalog locations [*Mercier de Lépinay et al.*, 2011], implying that the western
518 aftershock cluster is not onland but offshore. The Haiti-OBS catalog also confirms the
519 existence of a cluster of aftershocks more than 30 km west of the epicenter, 10-20 km
520 beyond the western end of coseismic slip in current source models.

521 In our kinematic source inversion based on teleseismic and geodetic data, slip is
522 allowed to spread over more than 45 km west of the hypocenter (= maximum source
523 duration \times maximum rupture speed = 15×3 km). However, in the resulting model the
524 slip remains confined to less than 22 km west of the epicenter. We obtained a similar
525 slip distribution in a finite fault inversion based only on the static data sets, InSAR
526 and GPS (Fig. S5 and S6) without constrain on the seismic moment, with comparable
527 fit to the data (Fig. S7 and S8) and with a seismic moment ($5.0 \cdot 10^{19}$ N.m) very close
528 to the GCMT solution ($4.7 \cdot 10^{19}$ N.m). Moreover, the InSAR fringes along the coast
529 (tracks A138a and A447d in Fig. S8) tend to rotate slightly perpendicular to western
530 end of our slip model, which suggests that there are no regions with larger slip further
531 west.

532 In summary, several arguments indicate that the geodetic data did not miss any
533 significant offshore deformation to the west. The spatial resolution of our regional
534 back-projection is too coarse in the source-to-array direction (NNW-SSE) to provide
535 an independent constrain on the distance between the coast and the second high-
536 frequency source. However, the position of this second source along the cross-range
537 direction (sub-parallel to the EPGF strike) is well constrained and is shown here to be
538 consistent with the western end of offshore slip inferred from static deformation. This
539 supports the idea that back-projection source imaging can identify the end tips of a
540 rupture.

541

542 **6.3. Advantage of regional array back-projection for earthquake source studies**

543

544 Our study of the 2010 Haiti earthquake demonstrates that back-projection of P-waves
545 recorded at regional distances can reveal the location of high frequency source

546 radiation that are not resolvable by finite source inversions at teleseismic distance,
547 which are now common practice. This new capability, improved by our analysis
548 procedure combining multitaper and MUSIC techniques, allows to study the rupture
549 of earthquakes with magnitude as low as 7 providing reliable spatial constraints. The
550 approach requires that the approximate strike of the fault is known, but this
551 information can usually be reliably extracted from the focal mechanism of the event.
552 This implies that regional back-projection will be critical in the study of earthquakes
553 when geodetic coverage is incomplete or local seismic networks are lacking, for
554 instance in subduction zones. Moreover, the procedure can be automated for rapid,
555 possibly real-time, earthquake analysis, combining the fault strike information from fast
556 moment tensor inversions. In that sense, the VNSN alone has a great potential for the
557 analysis of earthquake hazard for the whole Caribbean-North America plate
558 boundary. The high frequency aspects of the source process derived from array back-
559 projection are particularly useful to estimate potential damage in regions where the
560 vulnerable components of the building stock are dominated by low rise (short period)
561 buildings. In a more global perspective, the statistical analysis of the last 120 years of
562 earthquakes demonstrate that most of the devastating earthquakes occur in continental
563 interiors on previously unmapped faults [*England and Jackson, 2011*]. Thus, with the
564 ongoing development of regional seismic networks, many of those devastating
565 earthquakes will still happen at regional distance from networks, and require the
566 resolving power of regional back-projection.

567

568 **7. Conclusions**

569

570 In this study, we developed a back-projection source imaging technique for body
571 waves recorded by seismic arrays at regional distances (Pn waves). The technique
572 combines the Multiple Source Classification (MUSIC) method with multi-taper cross-
573 spectral estimation to achieve sharper source imaging than existing methods. This
574 technique allowed us to extract key aspects of the rupture process of the 2010 Haiti
575 earthquake from recordings by the Venezuela National Seismic Network (VNSN). In
576 particular, it provides the locations of high frequency source radiation. When
577 integrated with independent studies based on teleseismic and geodetic data of this
578 earthquake, our results indicate bilateral rupture at subshear speed, with a much
579 longer rupture segment towards the West. Prominent high frequency radiation
580 originates from rupture arrest phases at the tip of the main slip areas. Additional
581 rupture offshore is not supported by our analysis. Our results demonstrate how
582 regional array studies can contribute to the characterization of seismic sources in the
583 Caribbean region and elsewhere, particularly in offshore regions where local seismic
584 network or geodetic data coverage are not available, with potential application for
585 rapid earthquake response.

586

587

588 **Figure captions**

589 **Fig. 1 Haiti earthquake recorded by the Venezuela National Seismic Network.**

590 The triangles denote the 22 broadband stations of the VNSN that recorded the 2010
591 Haiti earthquake. The 13 stations shown in green were selected for our analysis based
592 on their mutual coherency. The red star indicates the epicenter (NEIC) of the Haiti
593 event. The inset shows vertical component seismograms filtered from 0.2 to 0.7 Hz,

594 aligned on their first P arrival and normalized by the standard deviation of their first
595 10 seconds.

596

597 **Fig. 2 Array response patterns**, of the VNSN (left) and USArray (right) back-
598 projected into the Haiti source region. plotted in map view. The color scale indicates
599 the power of the array response, normalized by its peak value. The yellow line
600 denotes the trace of the EPGF [*Calais et al.*, 2010]. The white dot is the location of
601 the NEIC epicenter.

602

603 **Fig. 3 Station selection based on array data correlation matrix** made of the
604 correlation coefficients between all pairs of stations (indicated by the color scale). The
605 stations are reordered by a clustering algorithm. The black box encompasses the most
606 mutually coherent subset of stations, which is used in our array analysis.

607

608 **Fig. 4 Dependence of MUSIC results on the assumed size N of the signal**
609 **subspace**, at 0.4 Hz with N=1 to 4 (a to d, respectively). The MUSIC pseudo-
610 spectrum as a function of along-strike position with respect to the hypocenter is
611 shown in color and normalized by its maximum in each time window. The time axis
612 is defined as the final time of the sliding windows of 30 s long sliding windows. The
613 horizontal axis is the distance along the fault with respect to the epicenter. The yellow
614 dashed lines and the black dots in figure b (N=2) show the bilateral rupture trend and
615 subevent locations.

616

617 **Fig. 5 Comparison of resolution between array processing techniques.** Two plane
618 waves, A and B, impinge on a linear array. The azimuth of A is fixed at 0 degrees
619 while the azimuth of B is varied from -10 to 0 degrees. Four array processing
620 techniques are considered: classical beamforming (a), cubic root stacking (b),
621 correlation stacking (c) and MUSIC (d). Each curve in the top plots shows the stack
622 (a-c) or pseudo-spectra (d) as a function of relative azimuth with respect to A for a
623 given azimuth separation between A and B (value indicated in the legend). The
624 bottom panels show the same quantities in color plots. The white dots mark the half-
625 width of the two largest maxima at a given azimuth of B. Correlation stacking
626 provides better resolution than beamforming and cubic root stacking. However,
627 MUSIC can resolve waves with azimuth separation as small as 3 degrees, achieving at
628 least twice higher resolution (minimum resolvable azimuthal separation) than the
629 other methods.

630

631 **Fig. 6 Synthetic test of array processing for the Haiti earthquake scenario.** The
632 earthquake is modeled by two point sources, the first one at the hypocenter and the
633 second one at a distance d westward along the fault trace. We applied cubic root
634 stacking (a), correlation stacking (b) and MUSIC (c). The left two columns are map
635 view back-projection images for $d = 30$ km (10s delay), in the time windows that start
636 at the beginning of the first ($t = 0$ s) and second ($t = 10$ s) source. The green asterisk
637 and red circle indicate the locations of the first and second source. The right two
638 columns show the projection of the images onto the fault trace, parallel to the source-
639 array direction, as a function of the distance d . The white lines indicate the locations
640 of the two sources. We considered moment ratios between first and second sources of
641 1:1 and 2:1. In the case of 2:1 source ratio, The cubic root stacking and correlation

642 stacking methods show resolution limit of about 35km, while the resolution limit of
643 MUSIC is about 25 km and the location error is approximately 5 km for each
644 subevent.

645 **Fig. 7 Synthetic and aftershock seismograms at the VNSN.** The top figure (a)
646 shows the vertical component of the synthetic seismograms of the mainshock filtered
647 from 0.2 to 0.7 Hz, assuming a point source with mechanism given by the CMT
648 solution. The bottom figure (b) shows the vertical recordings of a M5.9 aftershock
649 recorded at the VNSN with the same filtering applied.

650

651 **Fig. 8 Interference issues associated with the beamforming analysis.** Results of
652 two synthetic tests with a composite source comprising two subevents 15 km (top)
653 and 30 km (bottom) apart with Green's function computed with the Frequency-
654 Wavenumber method of [Zhu and Rivera, 2002]. Left: maps of beam amplitude
655 back-projected into the source region. Right: synthetic seismograms. The stations
656 deficient in low frequencies, due to interference effects, are colored in red.

657

658 **Fig. 9 Array analysis of the Haiti event.** Considering cubic root stacking (a),
659 correlation stacking (b) and MUSIC (c). The left two columns shows back-projection
660 images from 30s-long time windows ending at $t=10$ s and 30 s, respectively ($t=0$ s is
661 the arrival time). The yellow curve indicates the trace of the EPGF [Calais et al.,
662 2010]. The green line in the top-left plot is the major axis of the array response pattern
663 along which we project the array analysis onto the fault. The MUSIC analysis has
664 higher resolution and can clearly separate two asperities. The right column shows the

665 projection on the fault as a function of final time of the sliding window. The MUSIC
666 analysis of the largest aftershock (M5.9) is shown in panel (d).

667

668 **Fig. 10 Source time function from USArray data.** The triangles in the map show
669 the USArray stations that recorded the Haiti event. The triangles in blue are all
670 USArray stations that record the Haiti event. The stations in green have epicentral
671 distance greater than 30 degree and good signal to noise ratio, and were selected for
672 deconvolution. The red star is the epicenter of the Haiti earthquake. The red triangle is
673 station R23A. The inset shows (a) the recorded and synthetic seismograms, (b) the
674 source time function retrieved by non-negative least square (red) and Kikuchi-
675 Kanamori deconvolution techniques (green) at station R23A; and (c) source time
676 functions (green) from all selected USArray stations and the stacked source time
677 function (red). (d) histogram showing the delay between the two peaks from
678 bootstrapping the stacked source time functions and the best-fit Gaussian distribution
679 (red curve).

680

681 **Fig.11. Improved joint finite fault inversion of the Haiti event.** Surface projection
682 of the slip distribution inferred from the joint inversion of teleseismic, GPS and
683 InSAR data. The black and red arrows represent the recorded and model predicted
684 campaign GPS vectors respectively. The two orange dash lines mark the locations of
685 the subevents identified by MUSIC back-projection assuming our refined mainshock
686 epicenter as reference (red star). The black empty star is the USGS epicenter. The
687 small empty circles are the aftershock epicenters of the Haiti-OBS campaign. The
688 three inset maps show the location of the study area (top left), the measured and
689 predicted coastal uplift (top right) and the inverted source time function (lower right).

690 On the source time function plot, the blue and green curve show the contribution of
691 the Leogane and offshore fault segments respectively.

692

693

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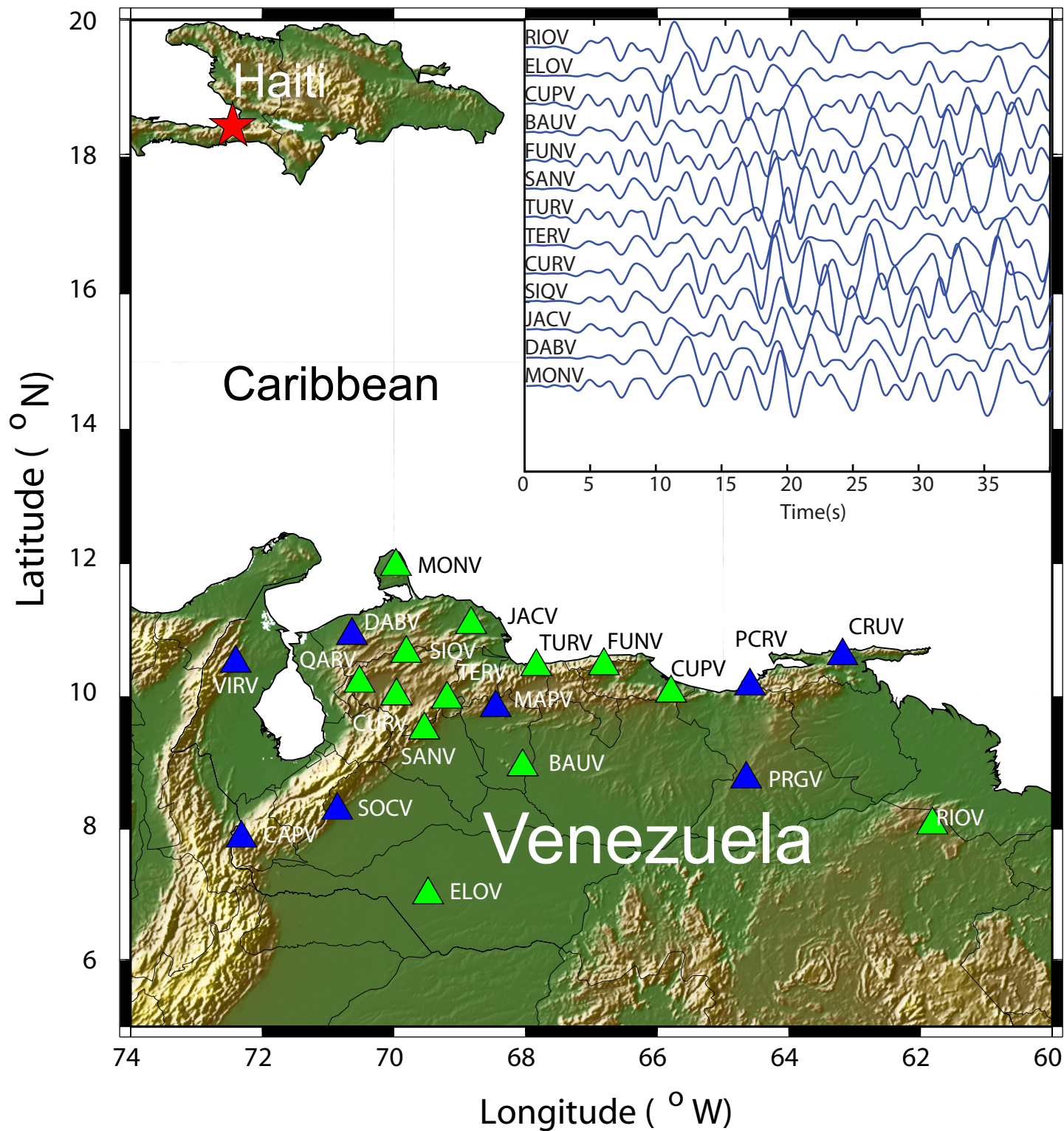
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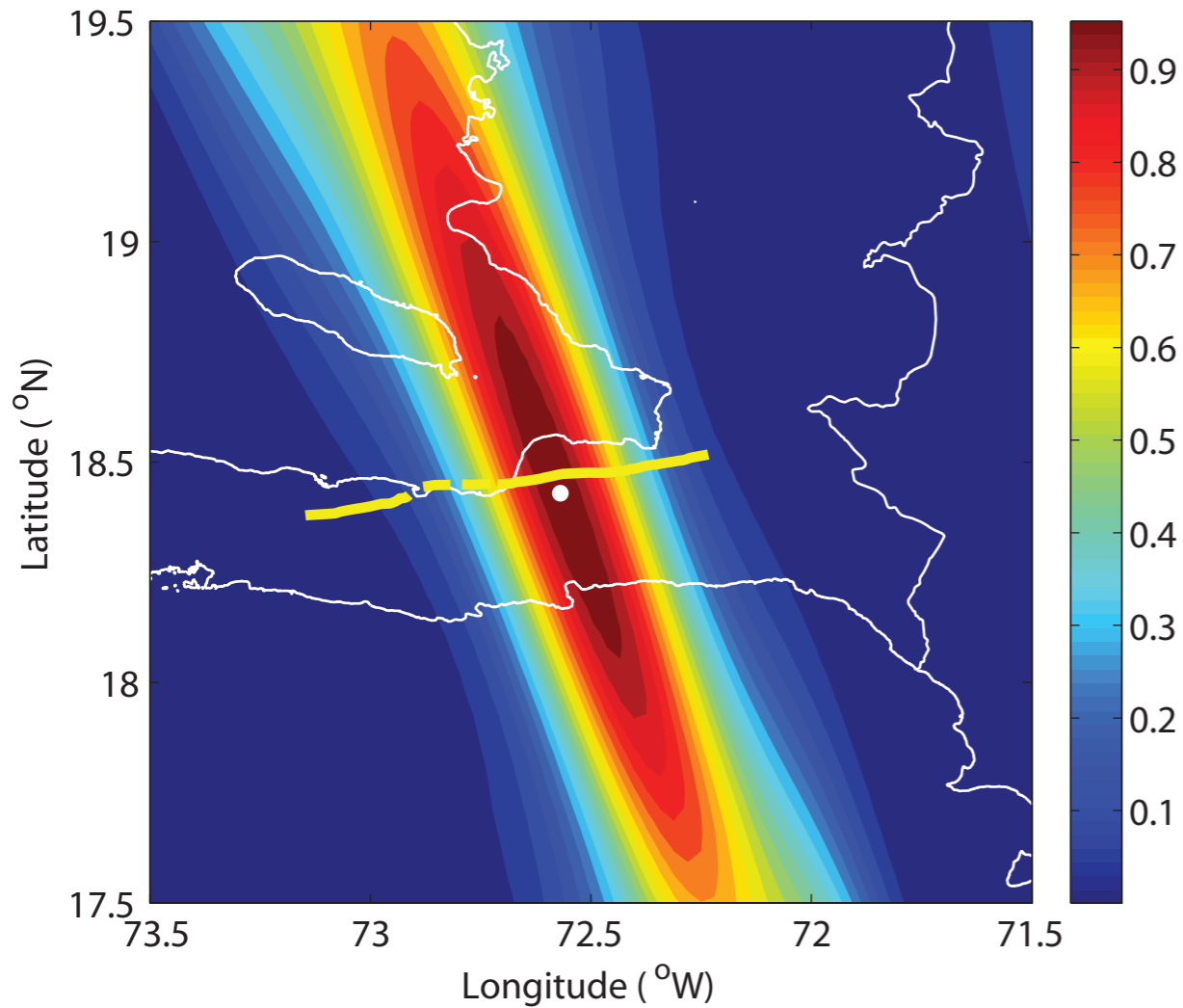
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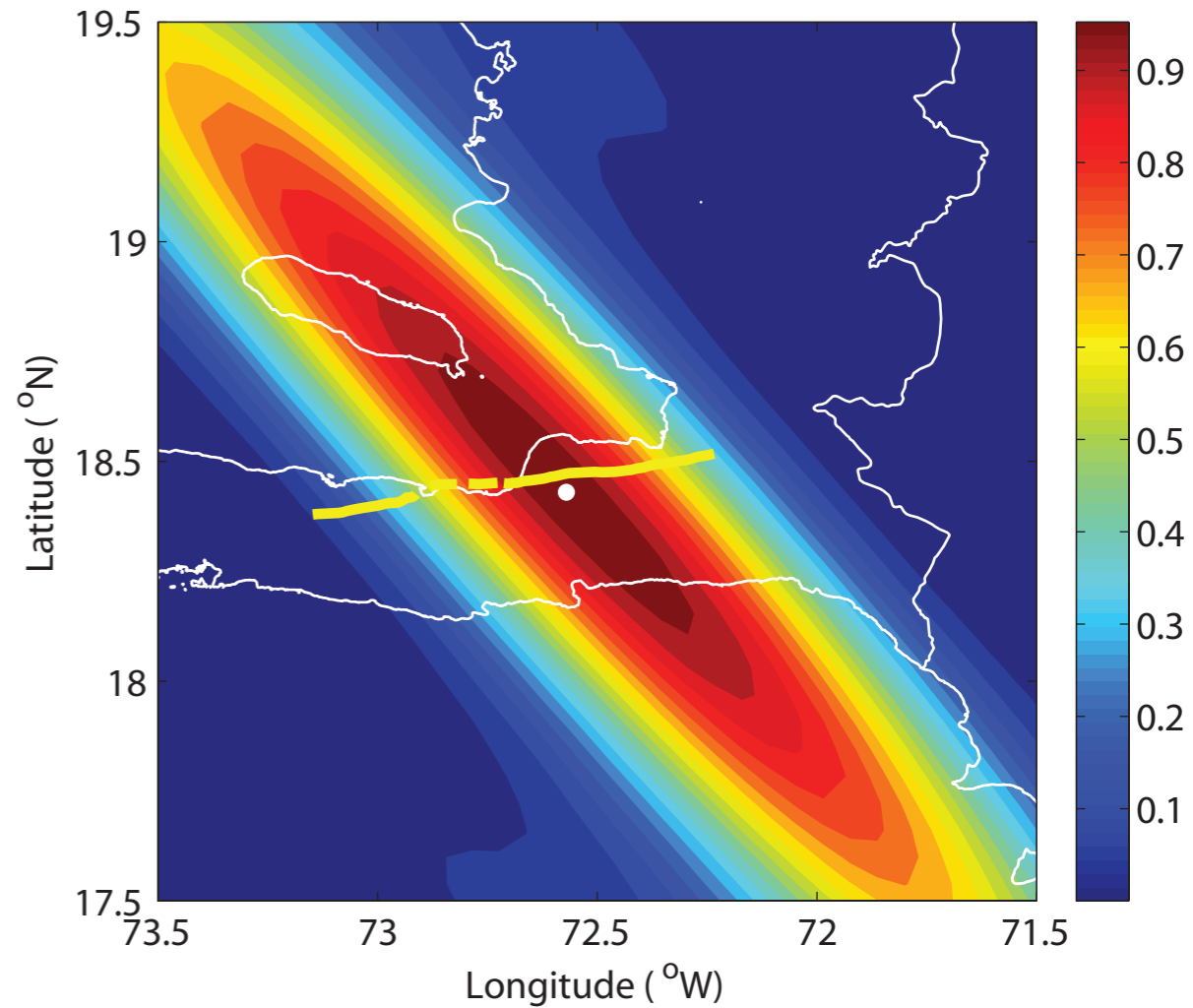
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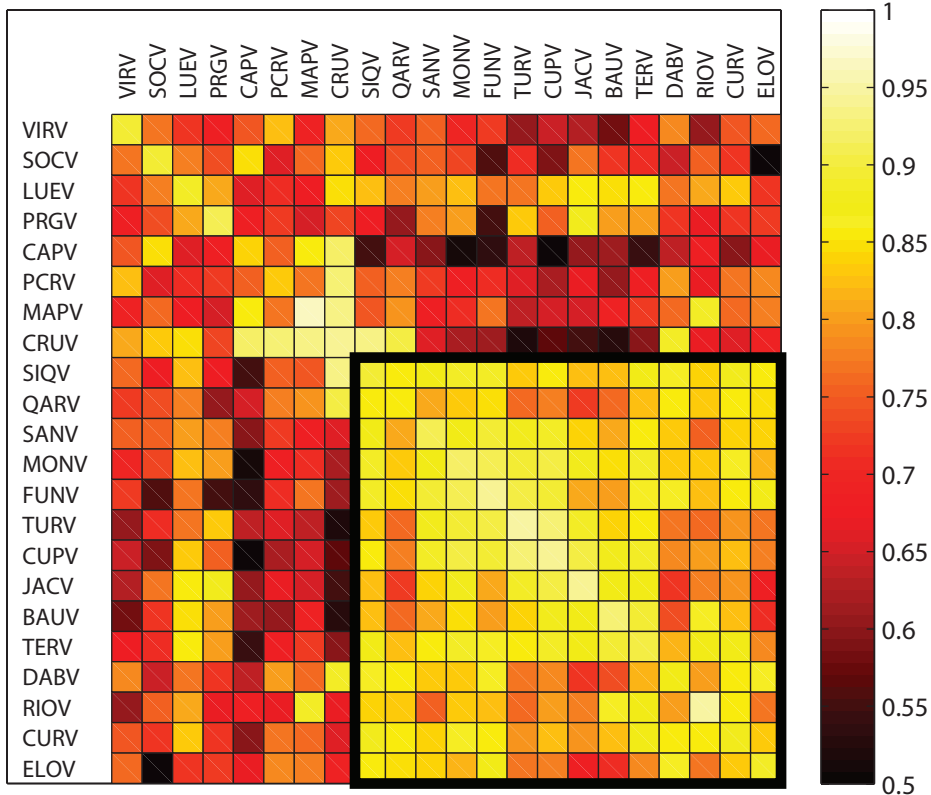


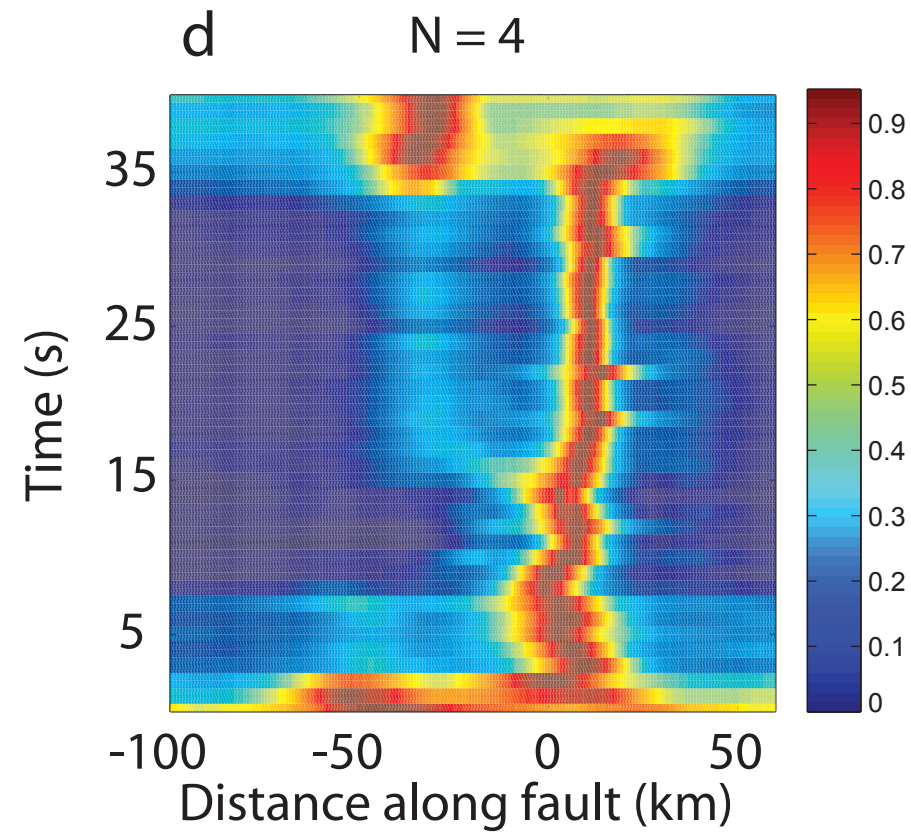
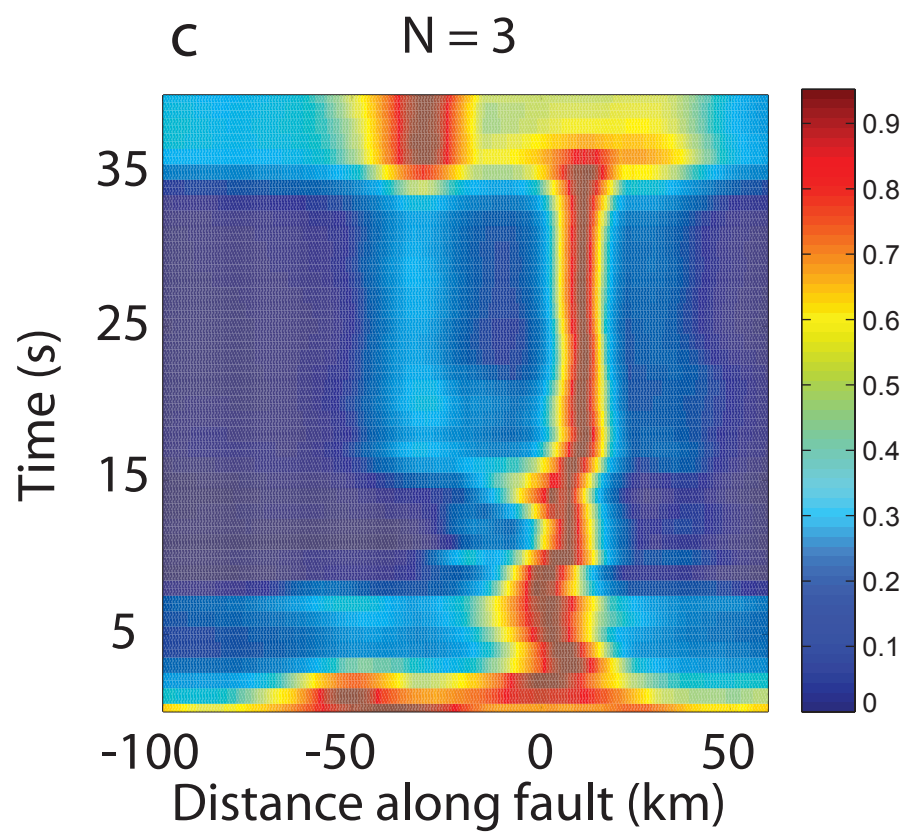
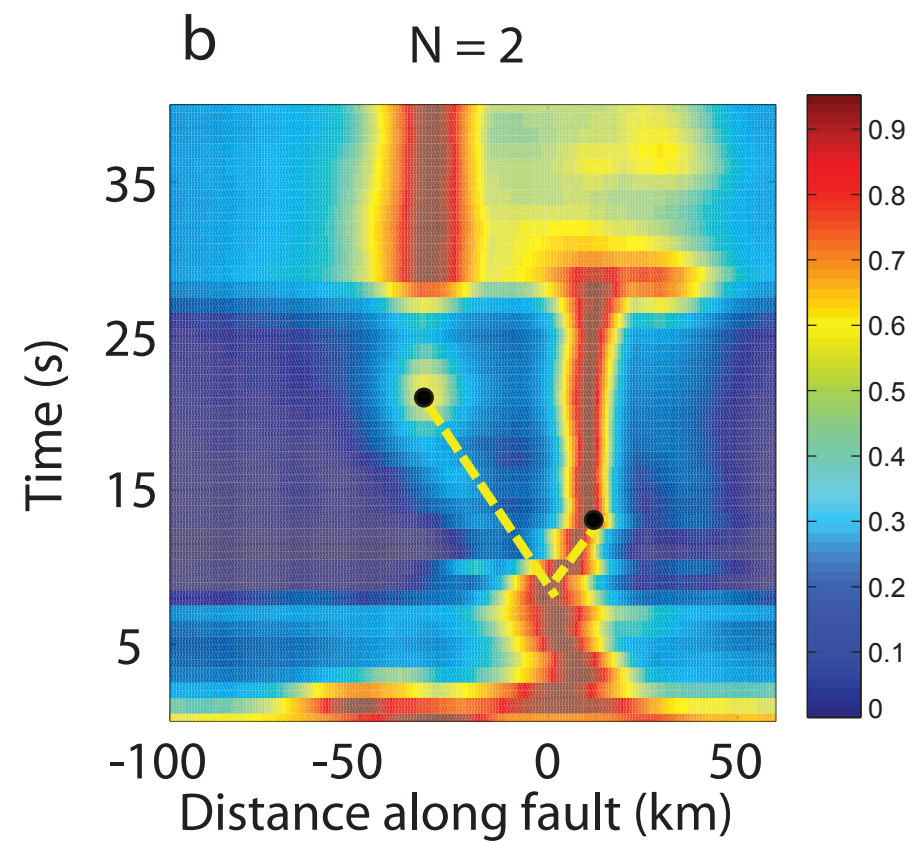
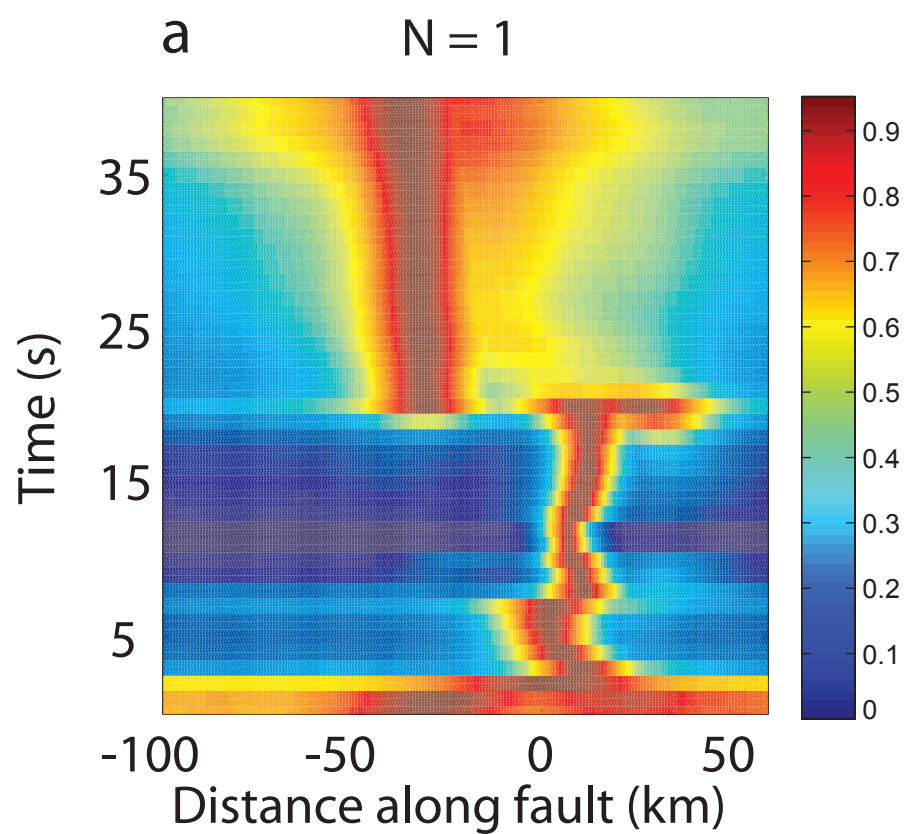
VNSN

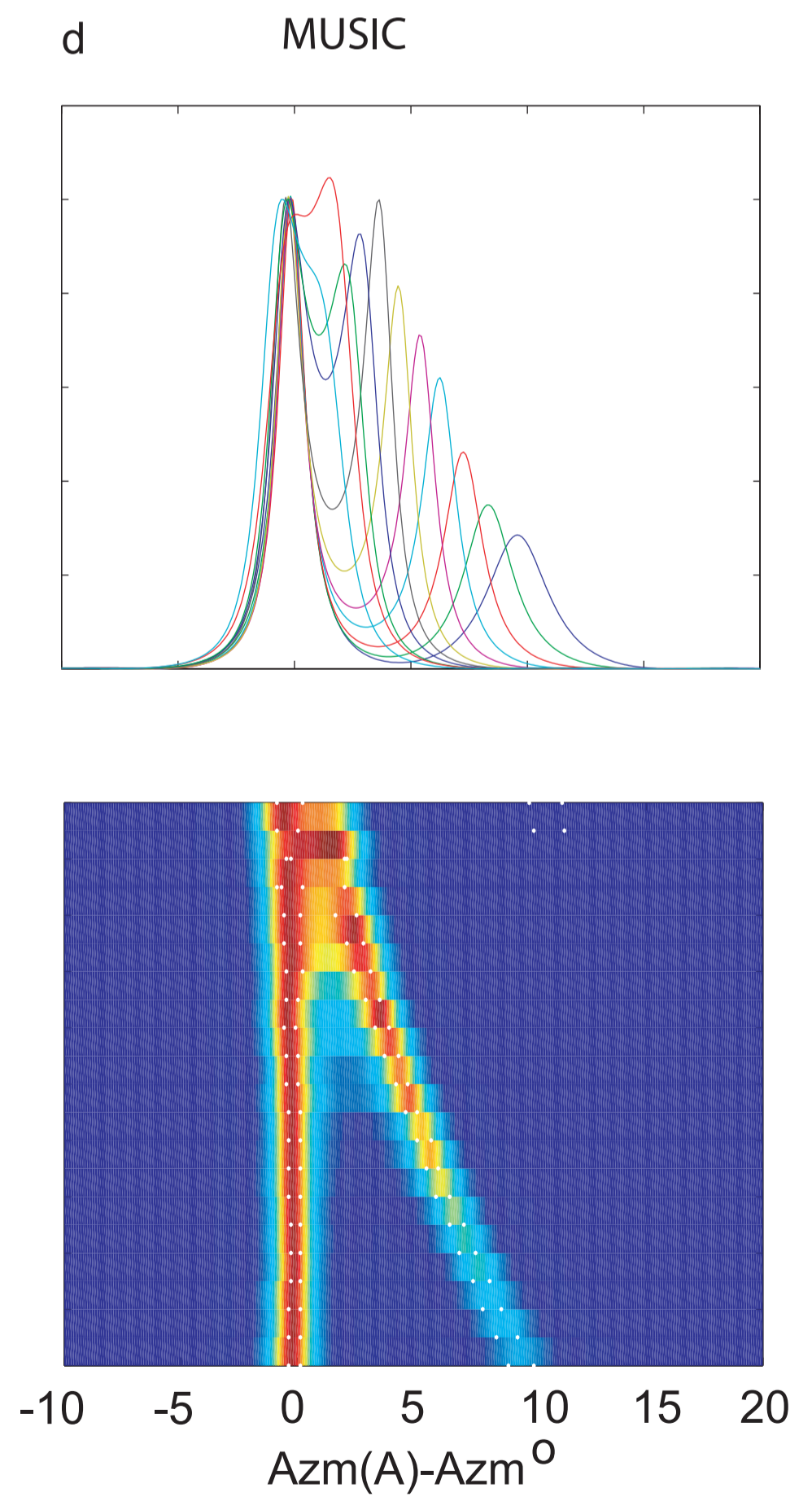
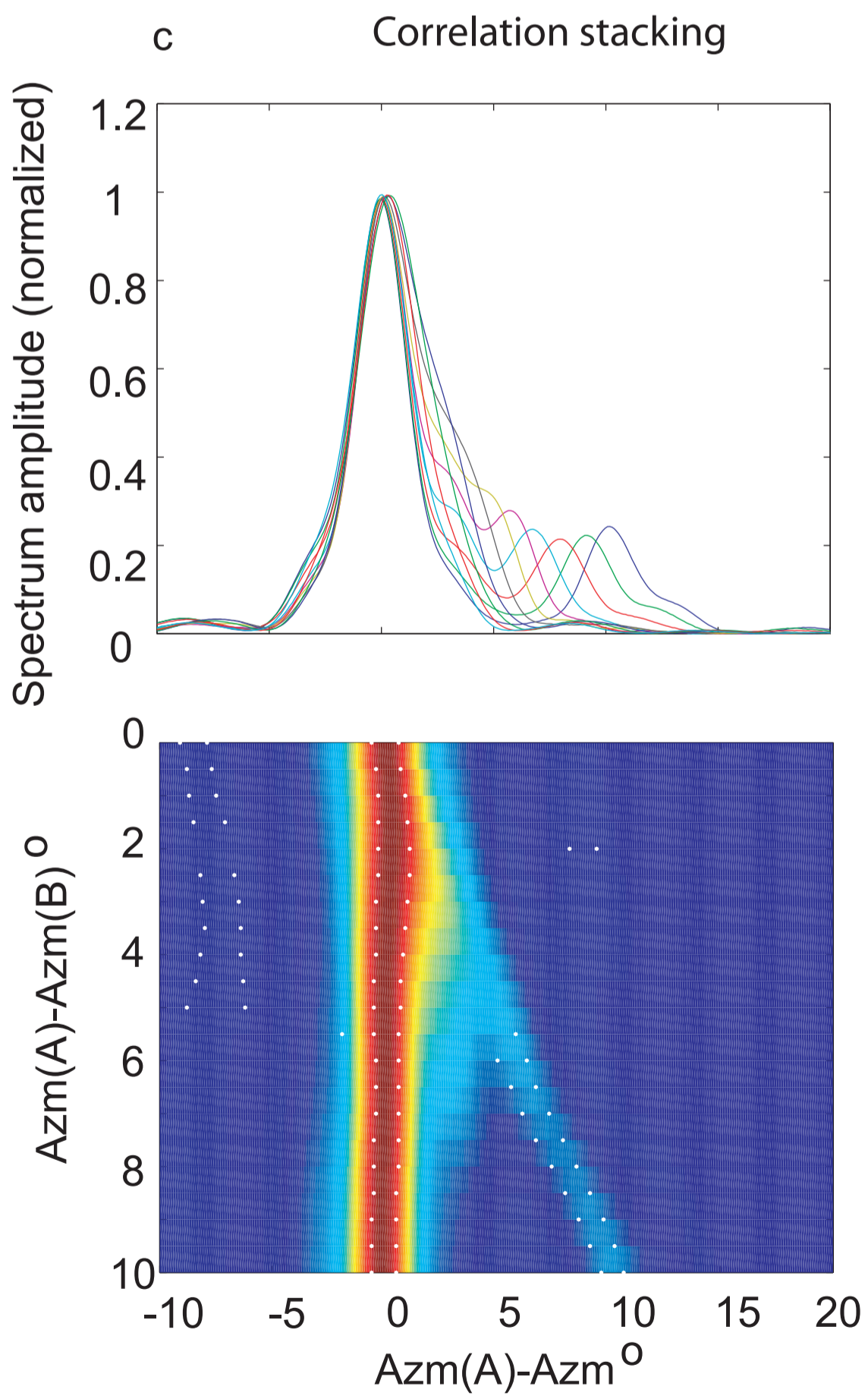
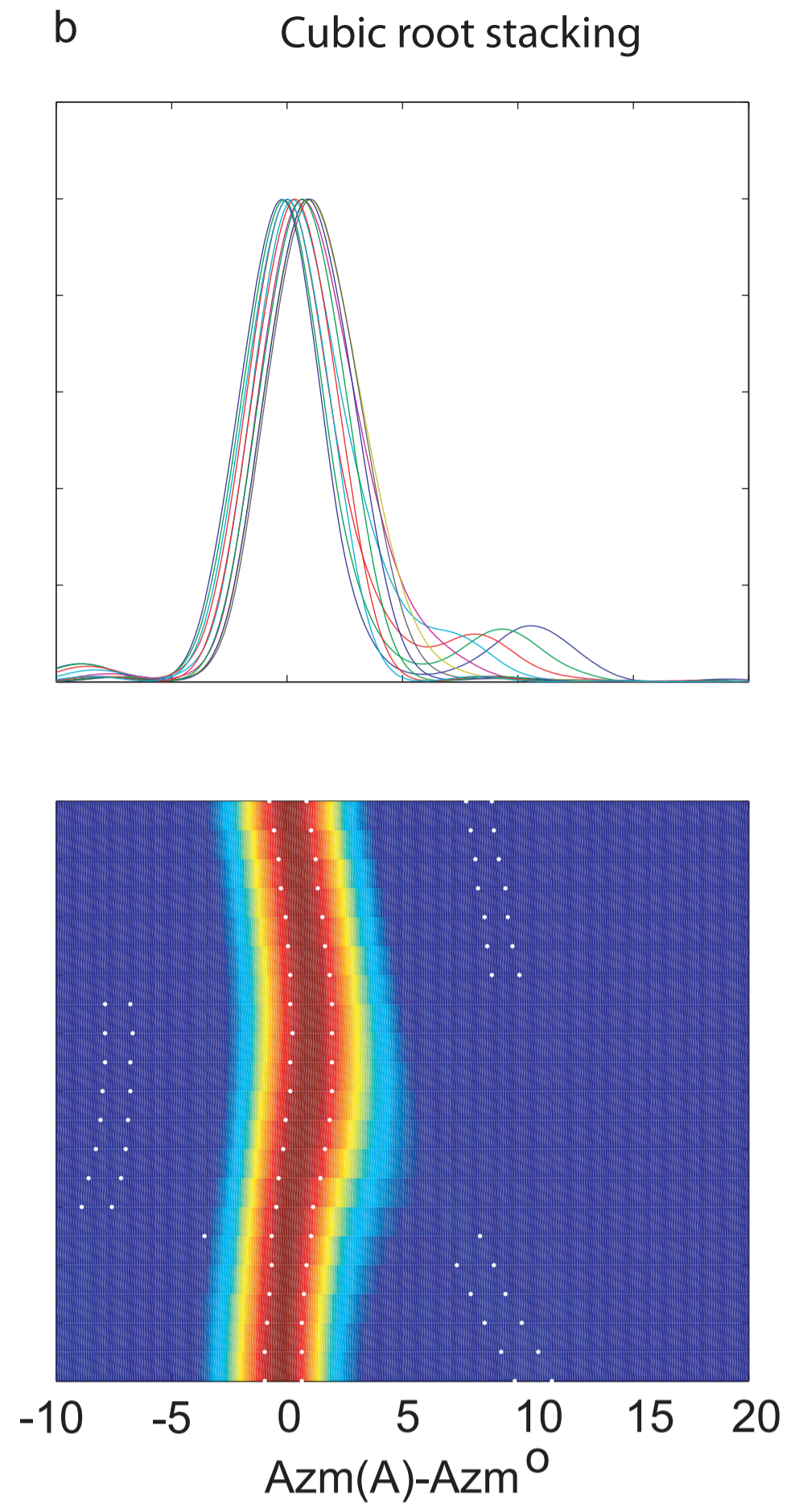
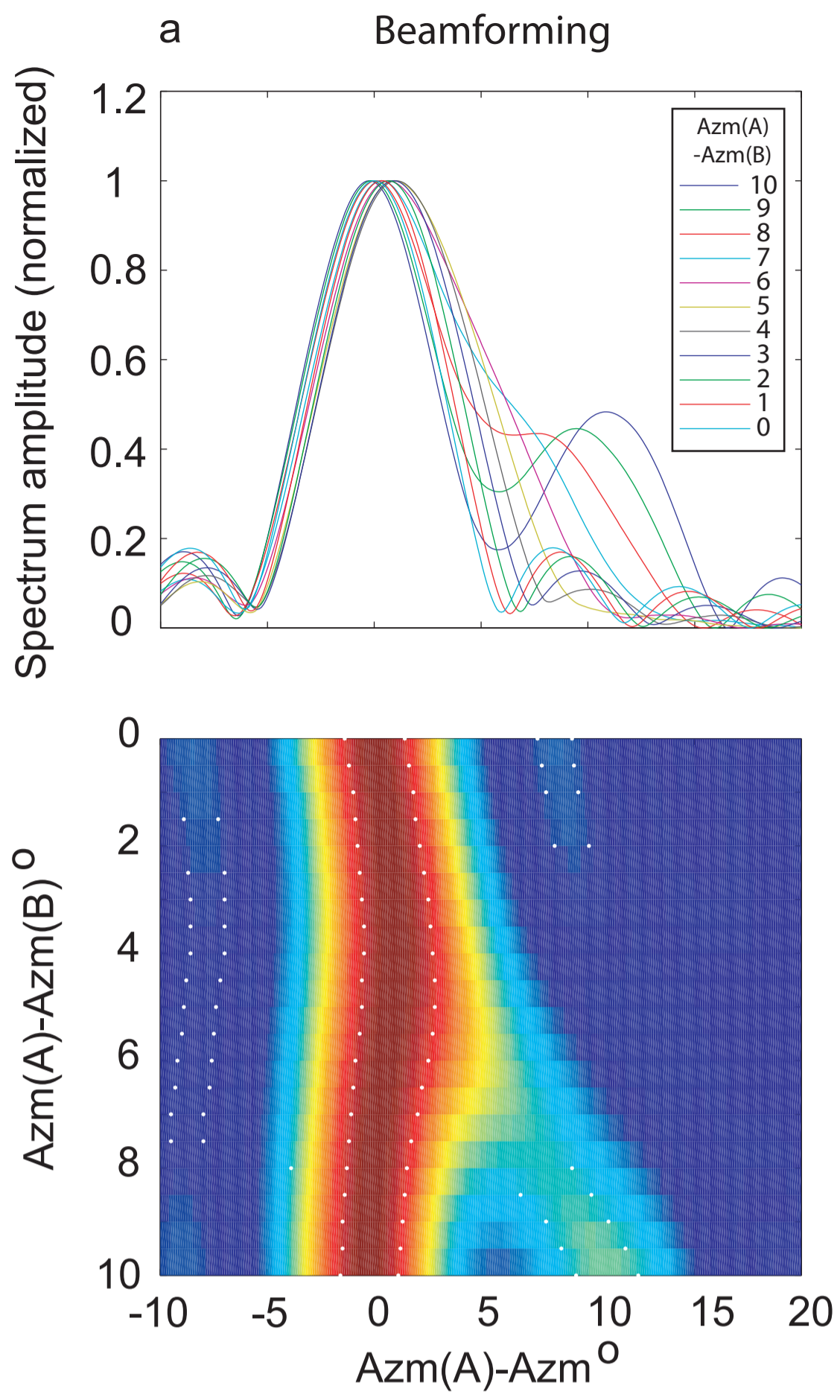


USArray

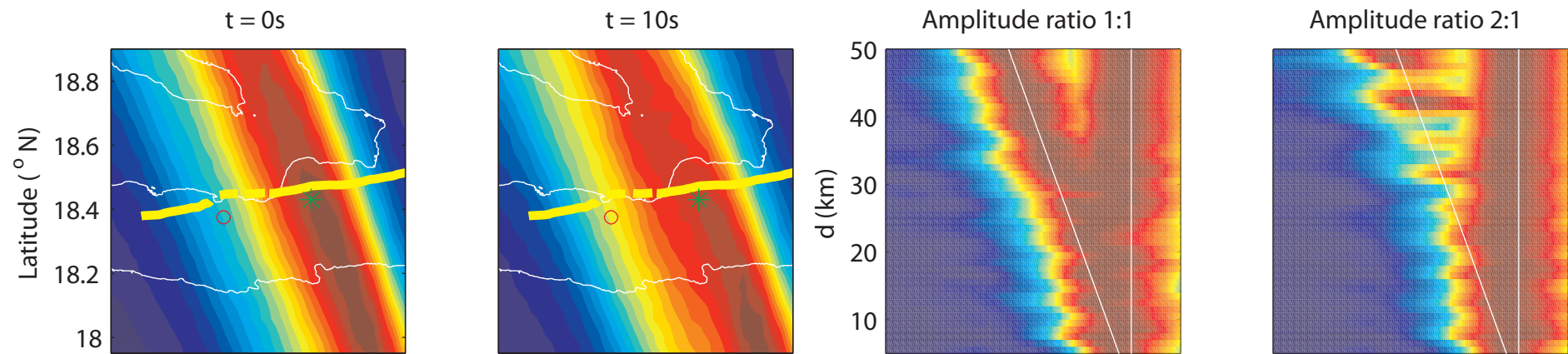




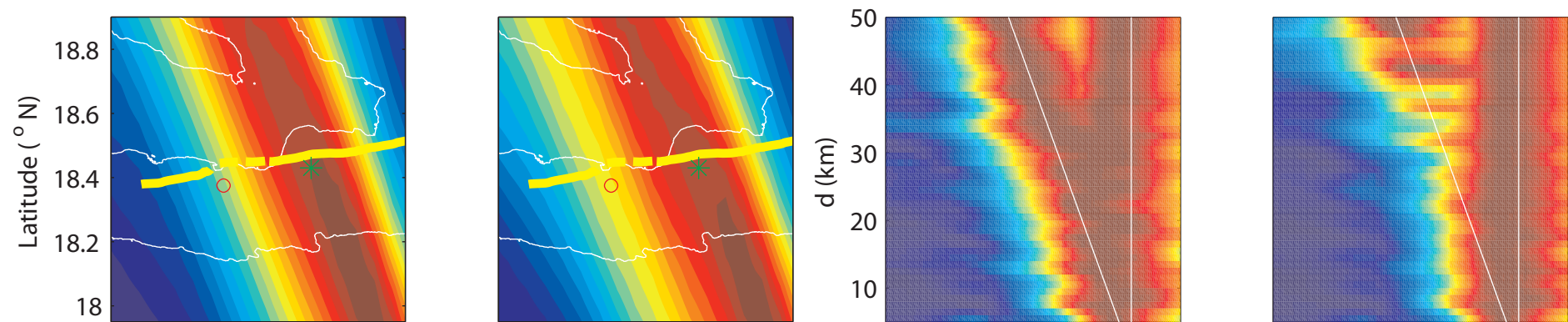




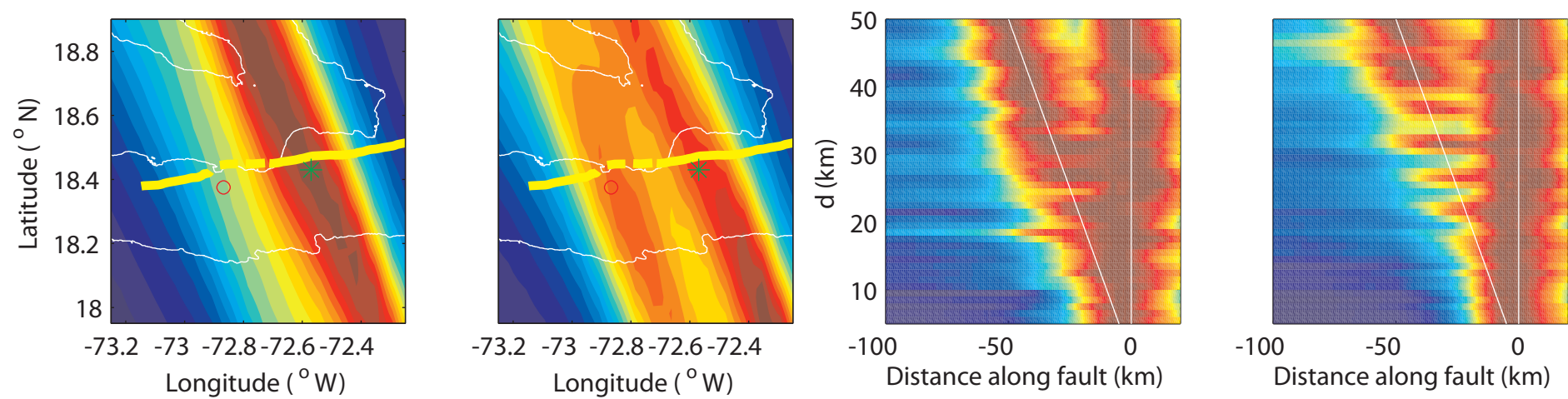
a Cubic root stacking



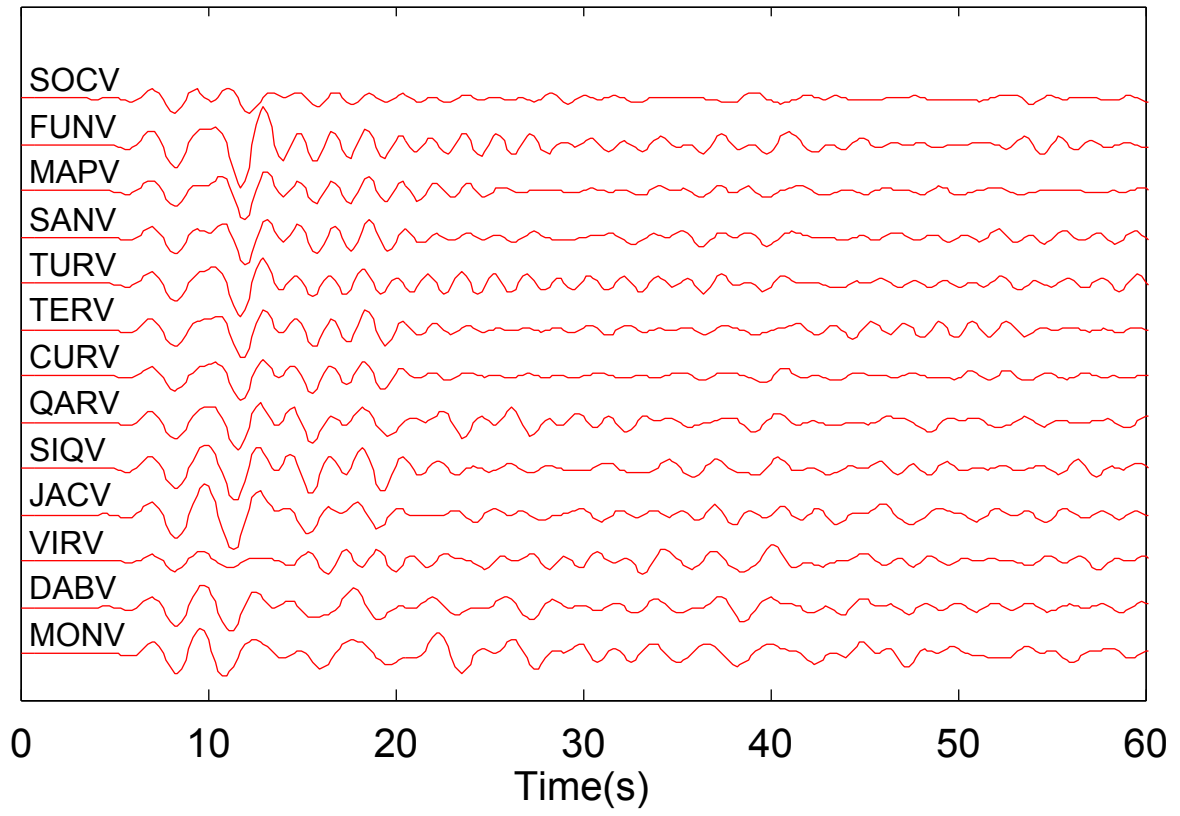
b Correlation stacking



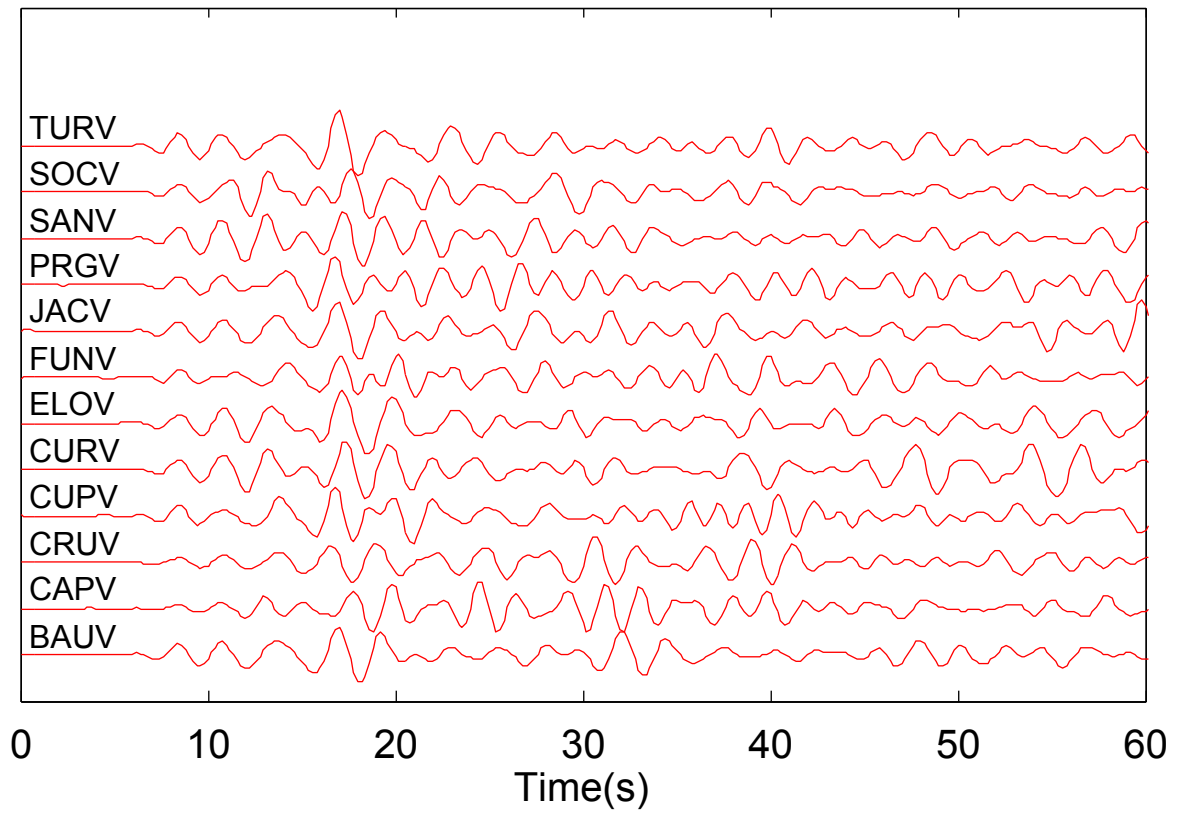
c MUSIC



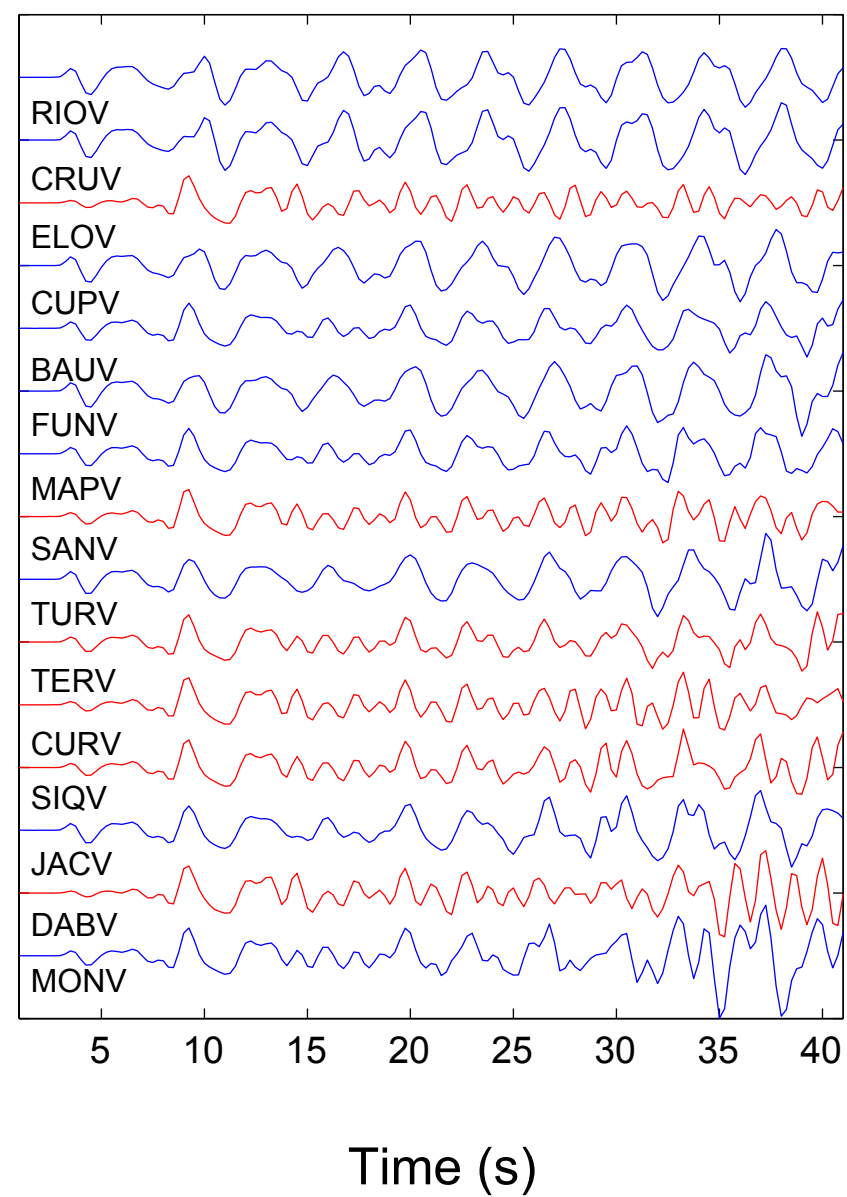
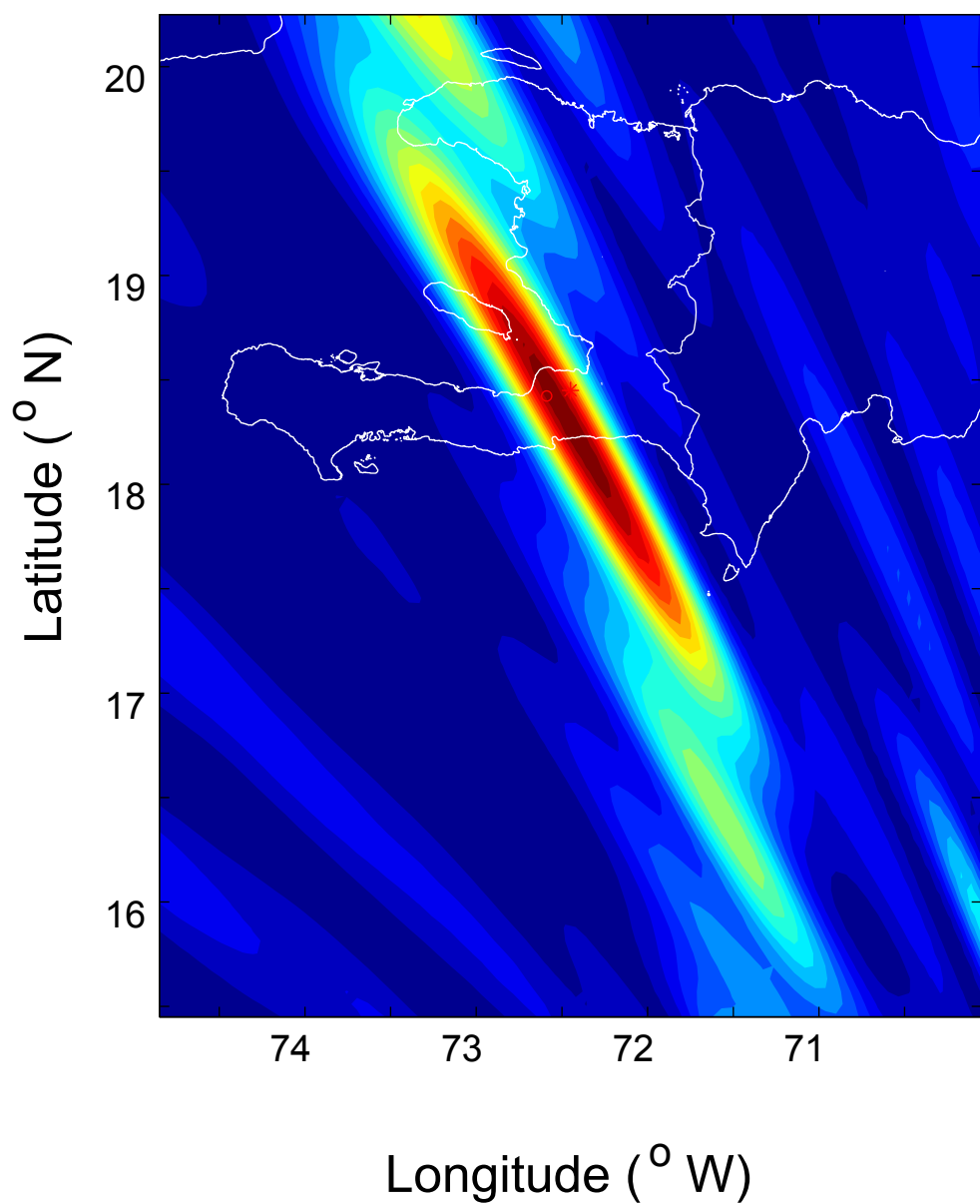
a Synthetic seismogram



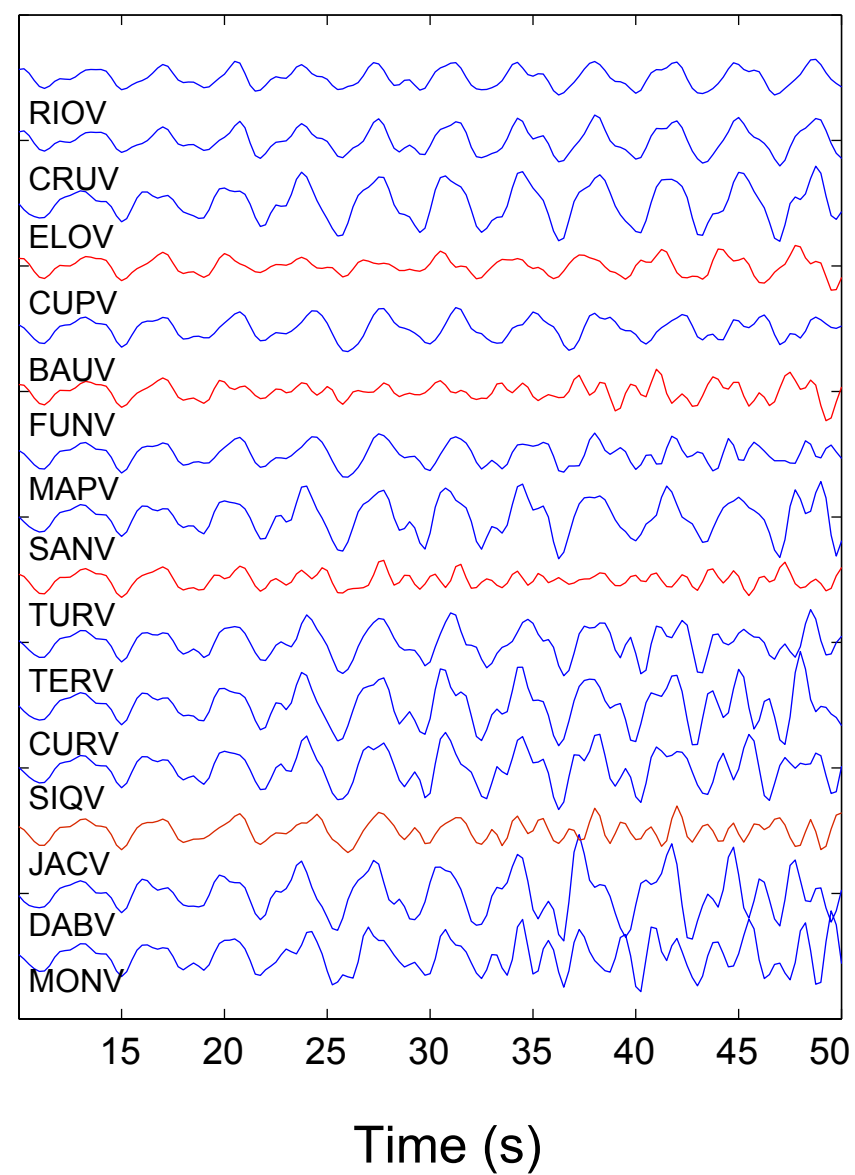
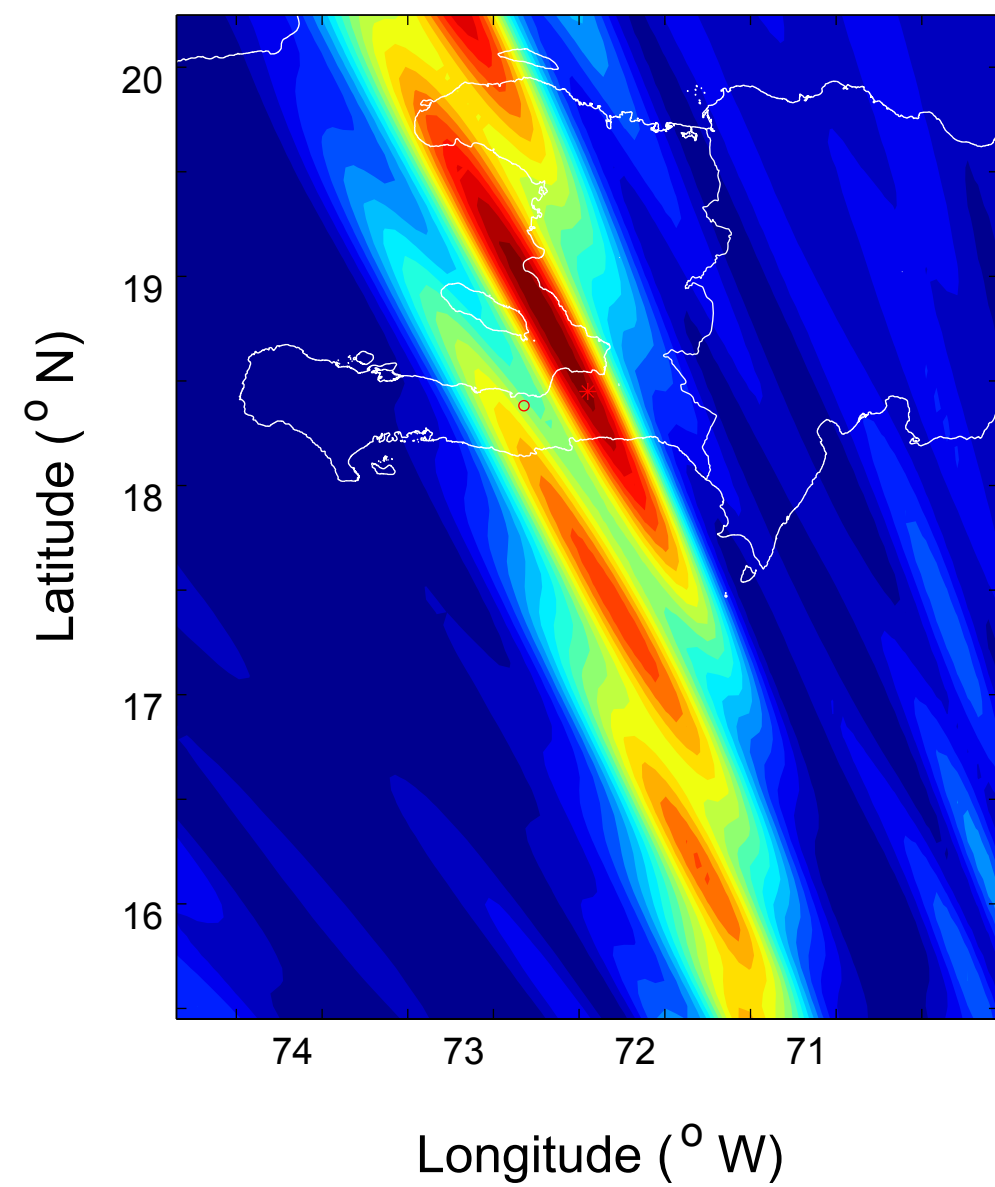
b Observed aftershock



Two sources 15km apart

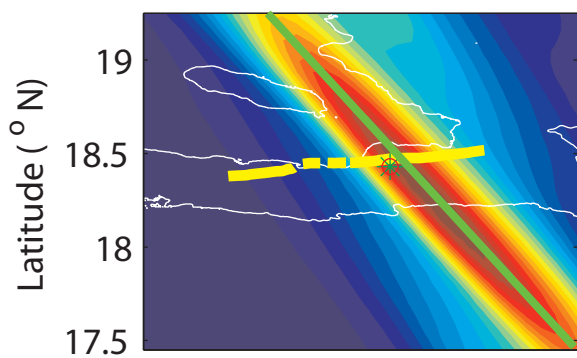


Two sources 30km apart

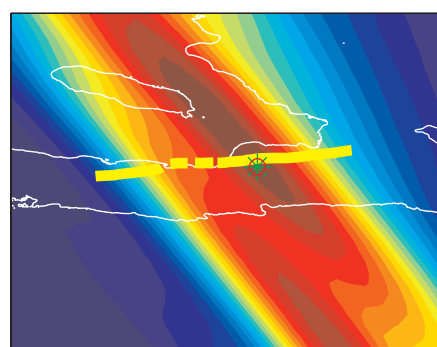


a Cubic root stacking

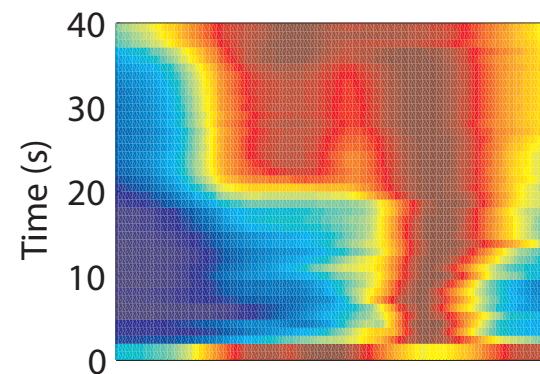
t = 15s



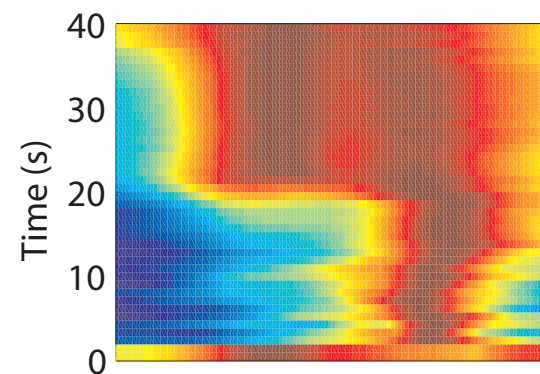
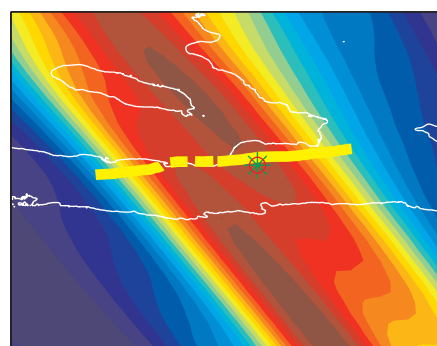
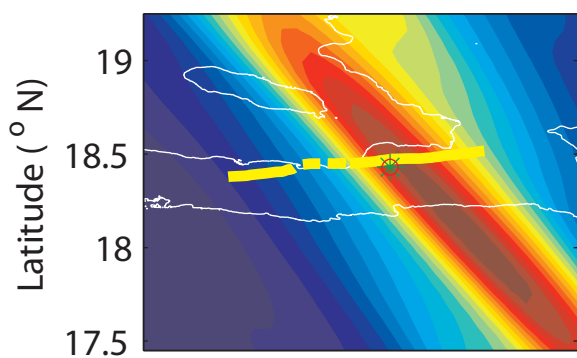
t = 30s



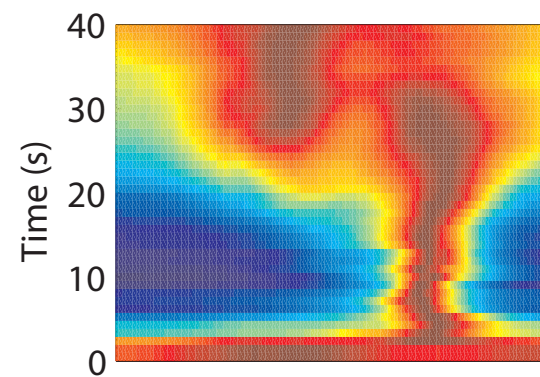
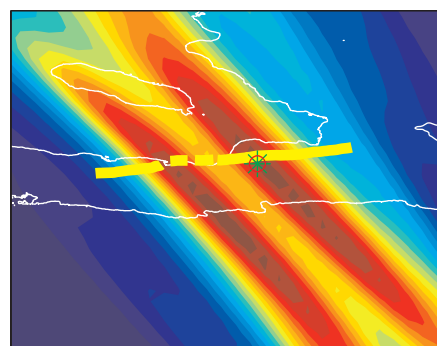
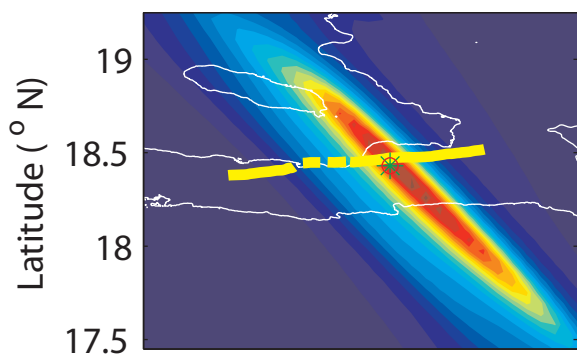
Evolution



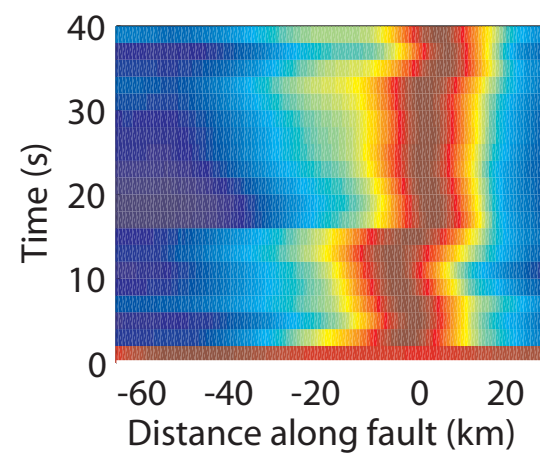
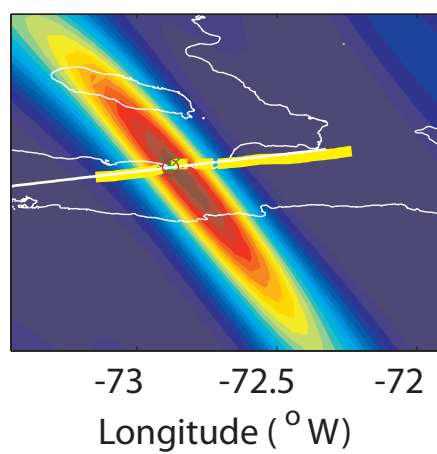
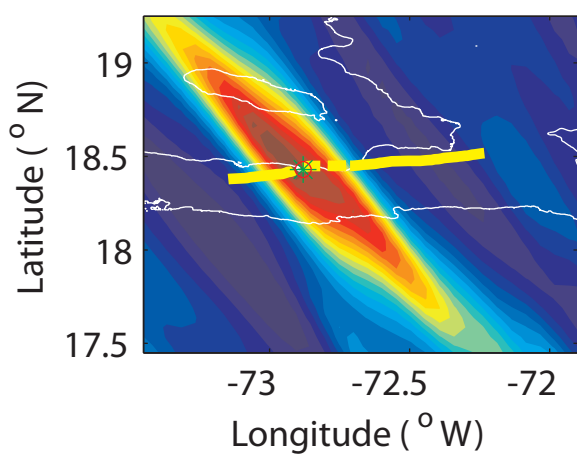
b Correlation stacking

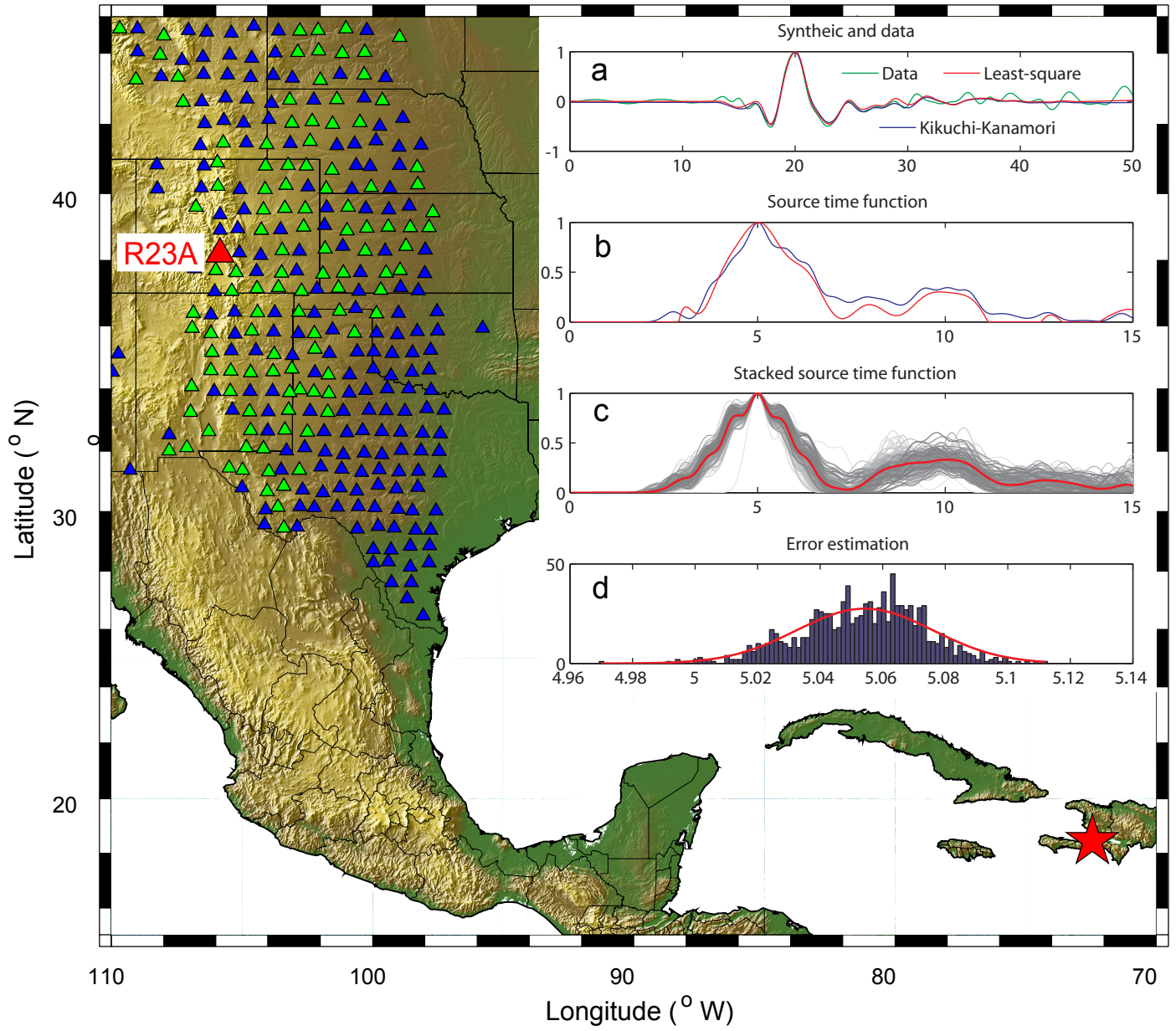


c MUSIC



d MUSIC (aftershock)





Kinematic model (InSAR+GPS+teleaseismic)

